## Course Roadmap

#### Rasterization Pipeline

#### **Core Concepts**

- Sampling
- Antialiasing
- Transforms

#### Geometric Modeling

#### **Core Concepts**

- Splines, Bezier Curves
- Topological Mesh Representations
- Subdivision, Geometry Processing

#### Lighting & Materials

#### **Core Concepts**

- Measuring Light
- Unbiased Integral Estimation
- Light Transport & Materials

#### Cameras & Imaging

Rasterization

**Transforms & Projection** 

**Texture Mapping** 

Visibility, Shading, Overall Pipeline

Intro to Geometry

**Curves and Surfaces** 

**Geometry Processing** 

Ray-Tracing & Acceleration

Radiometry & Photometry

Today

**Monte Carlo Integration** 

**Global Illumination & Path Tracing** 

**Material Modeling** 



#### Lecture 11:

# Measuring Light: Radiometry and Photometry

Computer Graphics and Imaging UC Berkeley CS184/284A

## Radiometry

Measurement system and units for illumination Measure the spatial properties of light

- New terms: Radiant flux, intensity, irradiance, radiance
   Perform lighting calculations in a physically correct manner
   Assumption: geometric optics model of light
  - Photons travel in straight lines, represented by rays

CS184/284A Ren Ng

## Light

#### Visible electromagnetic spectrum

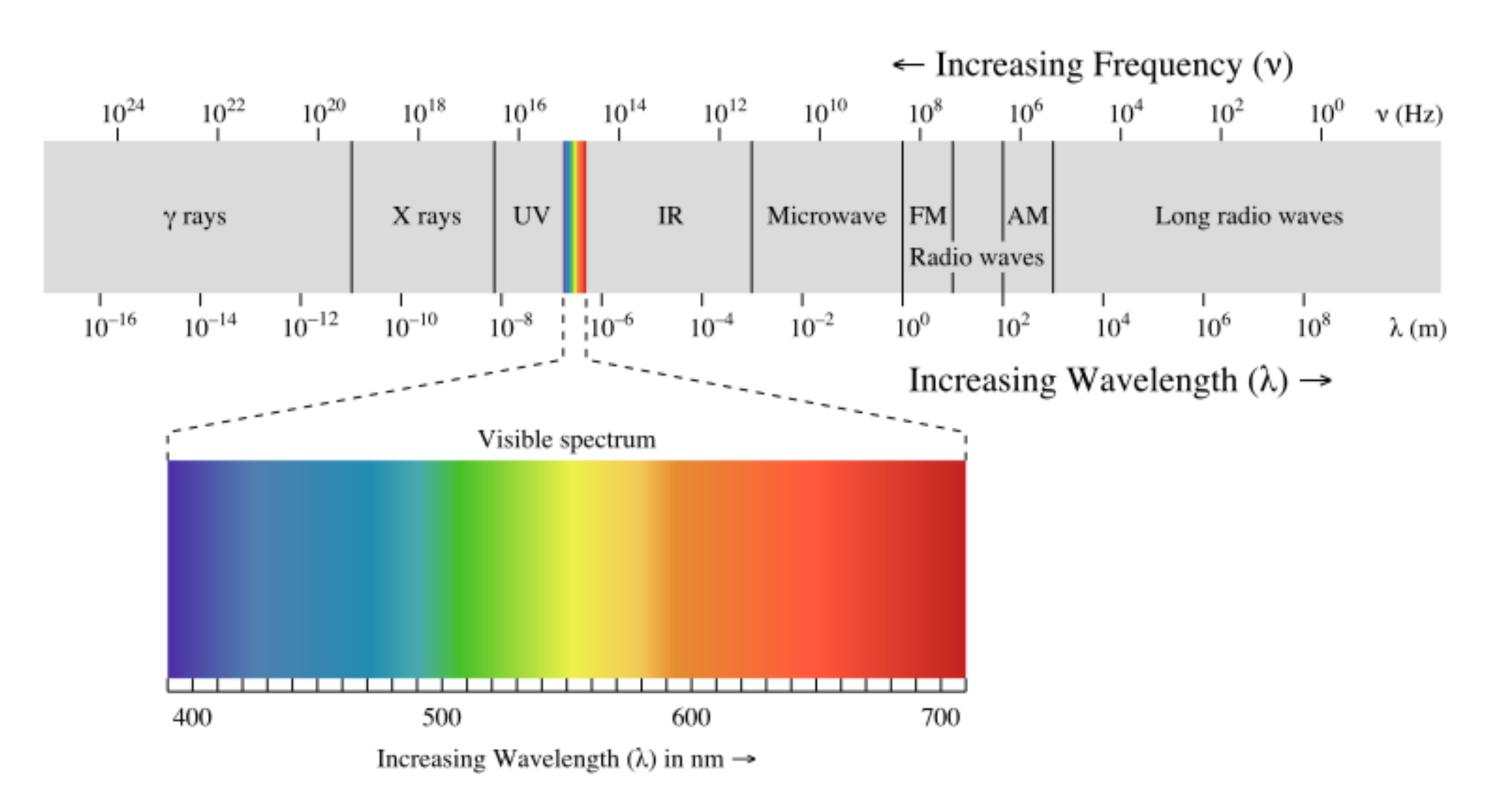


Image credit: Licensed under CC BY-SA 3.0 via Commons https://commons.wikimedia.org/wiki/File:EM\_spectrum.svg#/media/File:EM\_spectrum.svg

CS184/284A Ren Ng

## Lights: How Do They Work?



Cree 11 W LED light bulb (60W incandescent replacement)

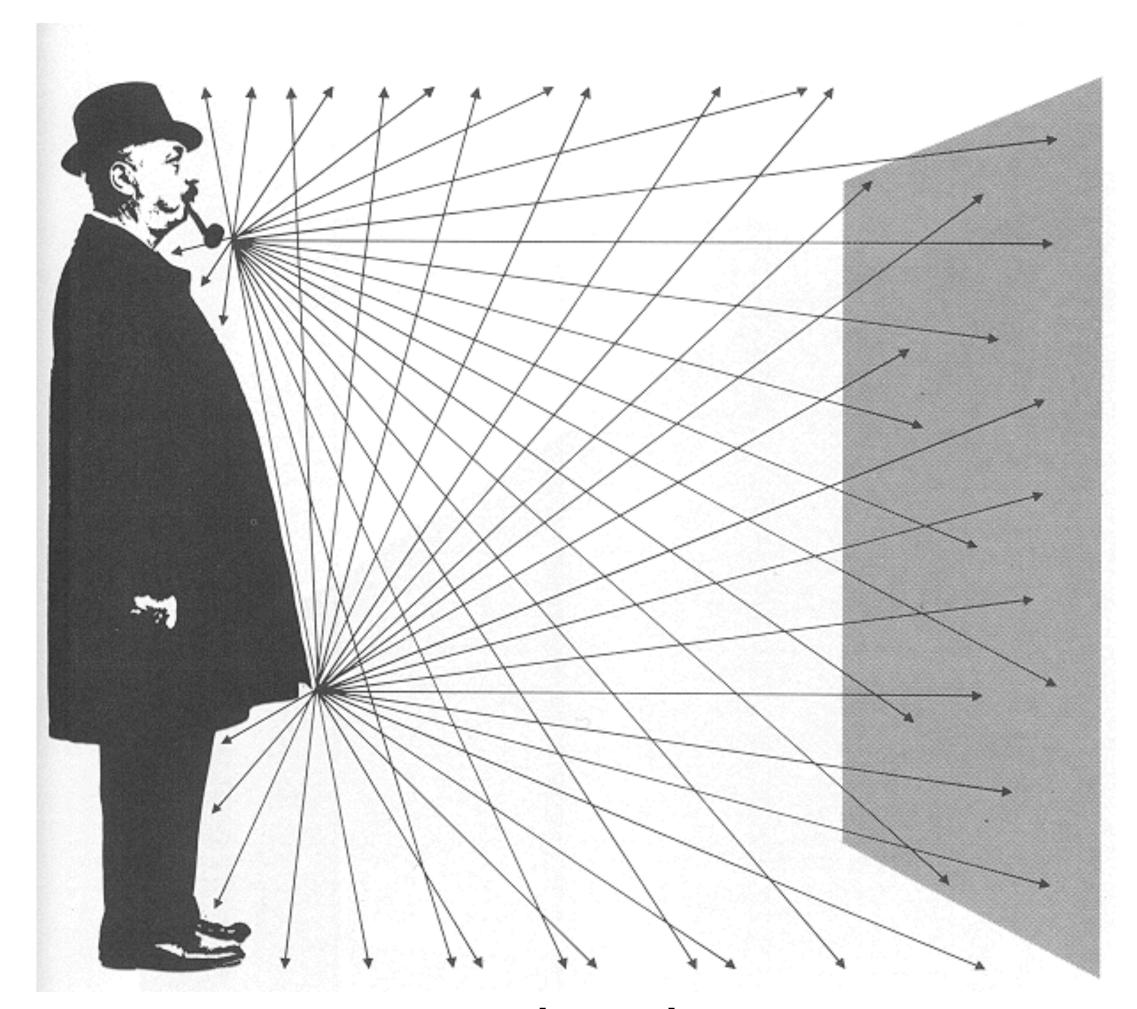
Physical process converts energy into photons

- Each photon carries a small amount of energy
   Over some amount of time, light consumes some amount of energy, Joules
- Some is turned into heat, some into photons
   Energy of photons hitting an object ~ exposure
  - Film, sensors, sunburn, solar panels, ...

Graphics: generally assume "steady state" flow

 Rate of energy consumption is constant, so flux (power) and energy are often interchangeable

#### Flux – How Fast Do Photons Flow Through a Sensor?



From London and Upton

CS184/284A Ren Ng

# Radiant Energy and Flux (Power)

## Radiant Energy and Flux (Power)

Definition: Radiant (luminous\*) energy is the energy of electromagnetic radiation. It is measured in units of joules, and denoted by the symbol:

$$Q$$
 [J = Joule]

Definition: Radiant (luminous\*) flux is the energy emitted, reflected, transmitted or received, per unit time.

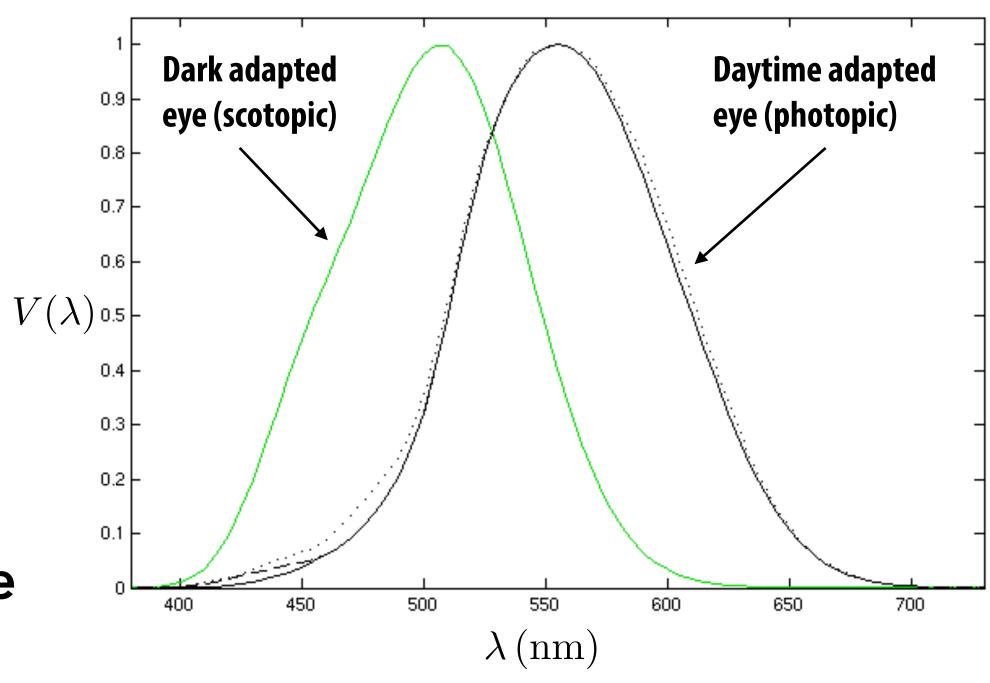
$$\Phi \equiv \frac{dQ}{dt} [W = Watt] [lm = lumen]^*$$

\* Definition slides will provide photometric terms in parentheses and give photometric units

CS184/284A

## Photometry

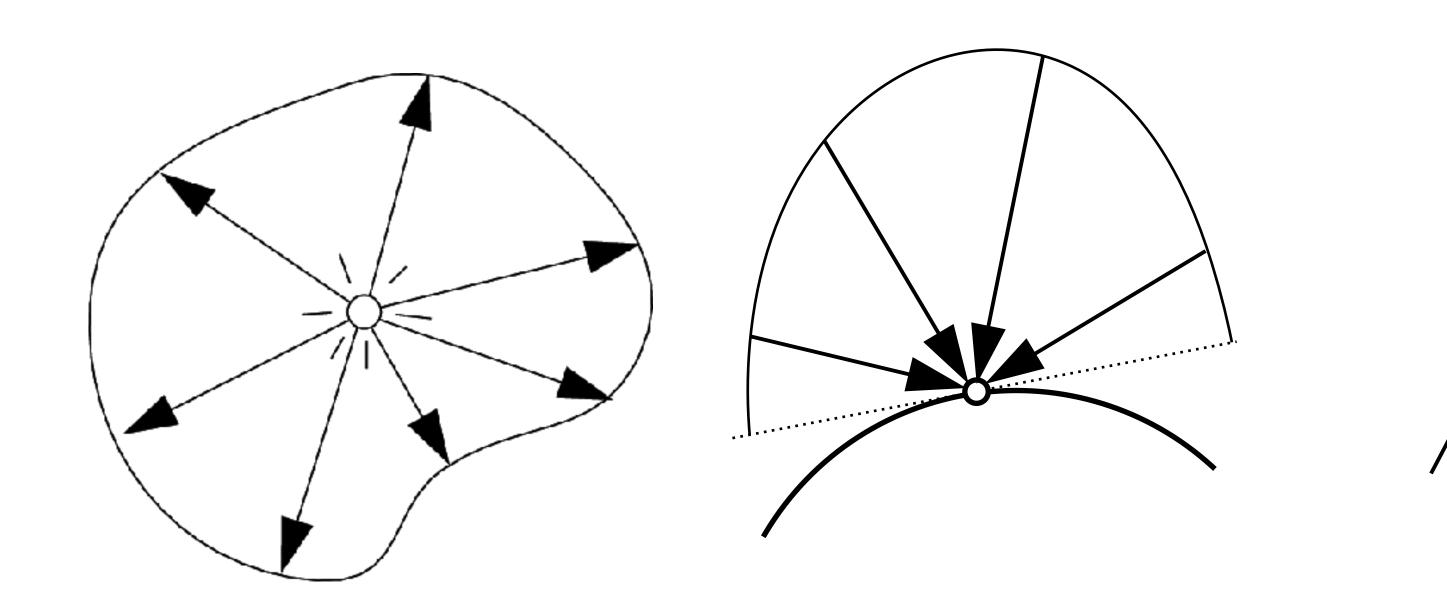
- All radiometric quantities have equivalents in photometry
- Photometry: accounts for response of human visual system
- E.g. Luminous flux  $\Phi_v$  is the photometric quantity that corresponds to radiant flux  $\Phi_e$ : integrate radiant flux over all wavelengths, weighted by eye's luminous efficiency curve  $V(\lambda)$



https://upload.wikimedia.org/wikipedia/commons/a/a0/Luminosity.png

$$\Phi_v = \int_0^\infty \Phi_e(\lambda) V(\lambda) \, d\lambda$$

#### Example Light Measurements of Interest



Light Emitted From A Source

"Radiant Intensity"

Light Falling
On A Surface

"Irradiance"

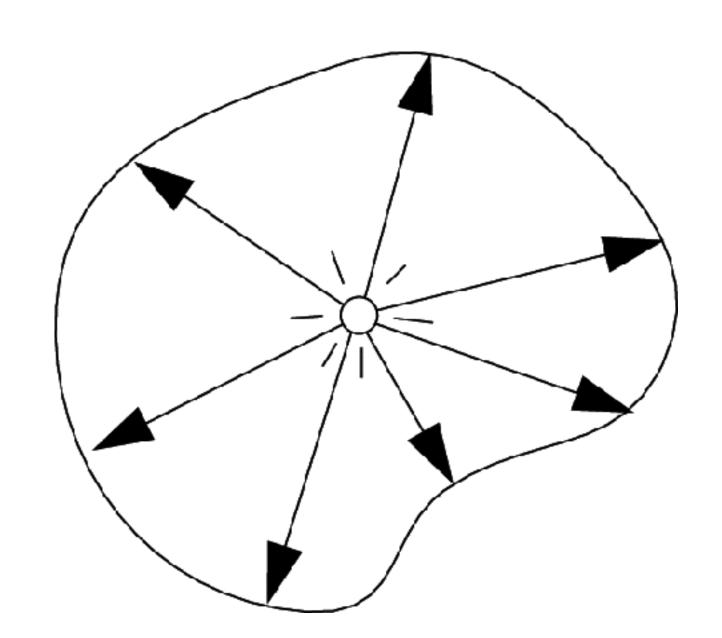
Light Traveling
Along A Ray

"Radiance"

# Radiant Intensity

## Radiant Intensity

Definition: The radiant (luminous) intensity is the power per unit solid angle emitted by a point light source.



$$I(\omega) \equiv \frac{\mathrm{d}\Phi}{\mathrm{d}\omega}$$

$$\left[\frac{W}{sr}\right] \left[\frac{lm}{sr} = cd = candela\right]$$

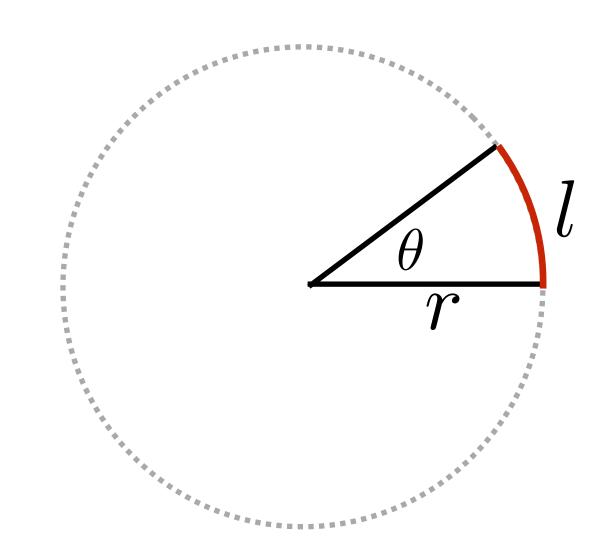
The candela is one of the seven SI base units.

## Angles and Solid Angles

Angle: ratio of subtended arc length on circle to radius

$$ullet \theta = rac{l}{r}$$

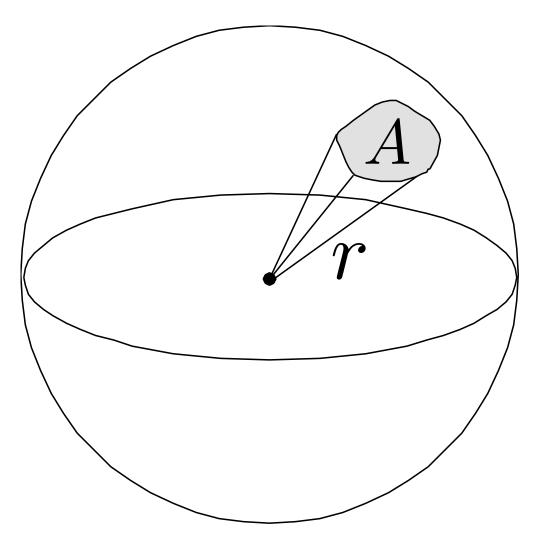
• Circle has  $2\pi$  radians



Solid angle: ratio of subtended area on sphere to radius squared

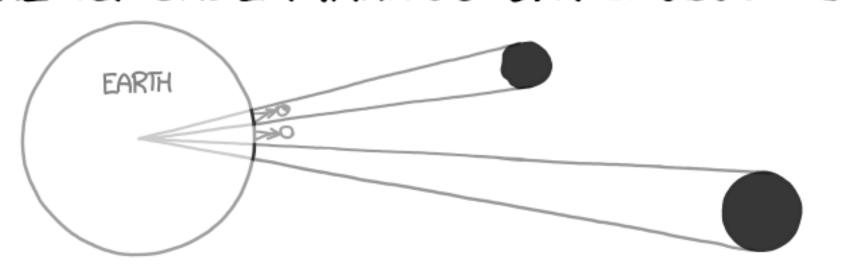
$$\bullet \ \Omega = \frac{A}{r^2}$$

• Sphere has  $4\pi$  steradians

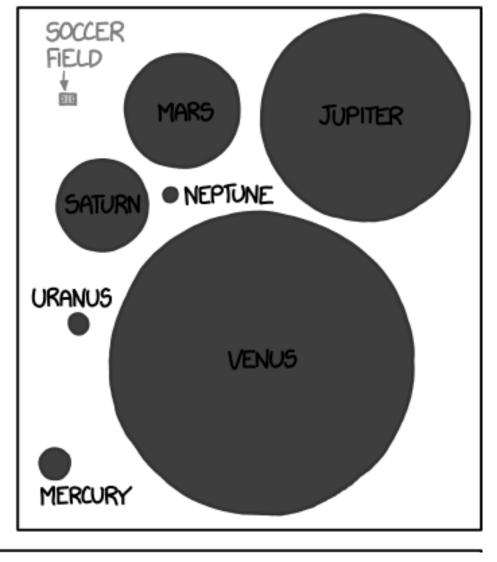


## Solid Angles in Practice

THE SIZE OF THE PART OF EARTH'S SURFACE DIRECTLY UNDER VARIOUS SPACE OBJECTS





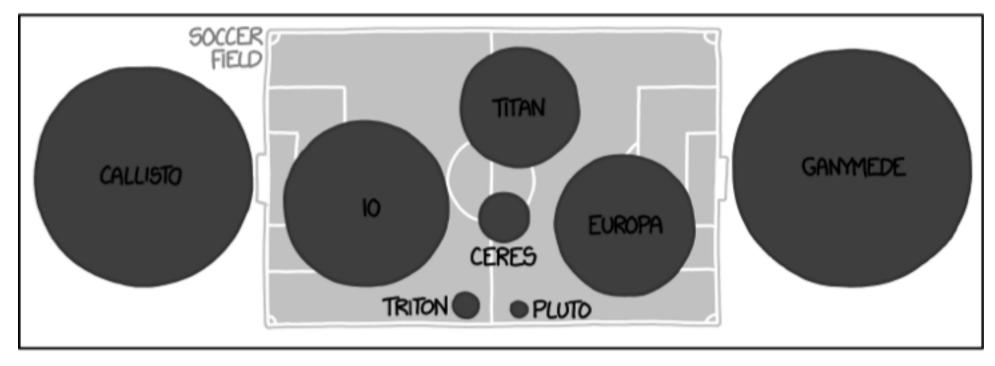


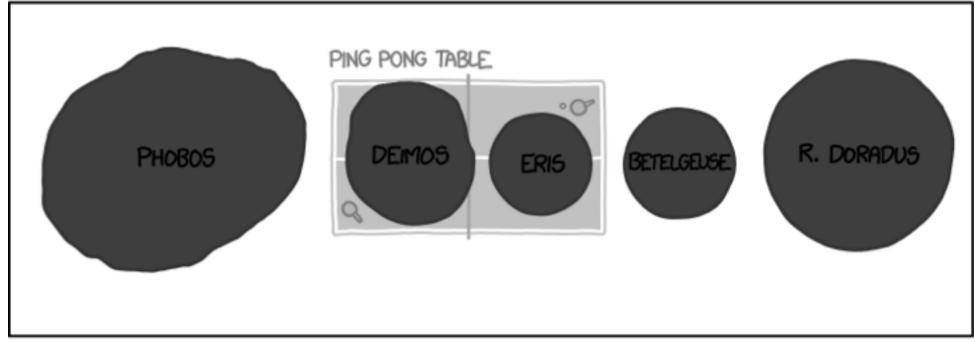
- Sun and moon both subtend ~60µ sr as seen from earth
- Surface area of earth:
   ~510M km²
- Projected area:

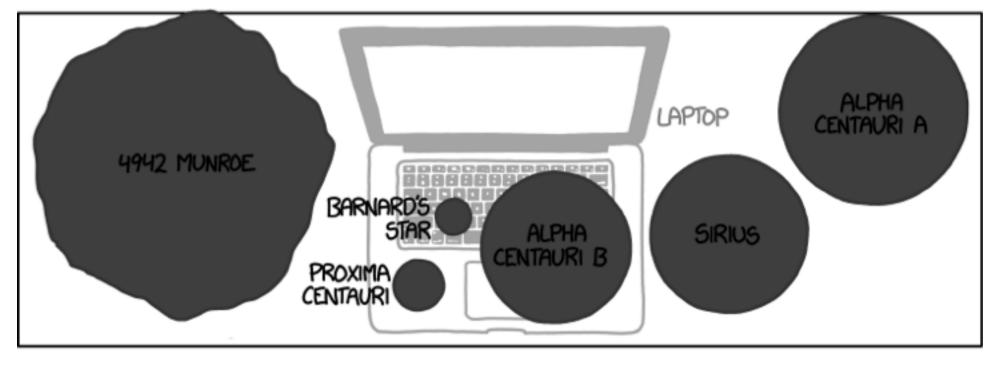
$$510 \text{Mkm}^2 \frac{60 \mu \text{sr}}{4 \pi \text{sr}} = 510 \frac{15}{\pi}$$
$$\approx 2400 \text{km}^2$$

http://xkcd.com/1276/

## Solid Angles in Practice

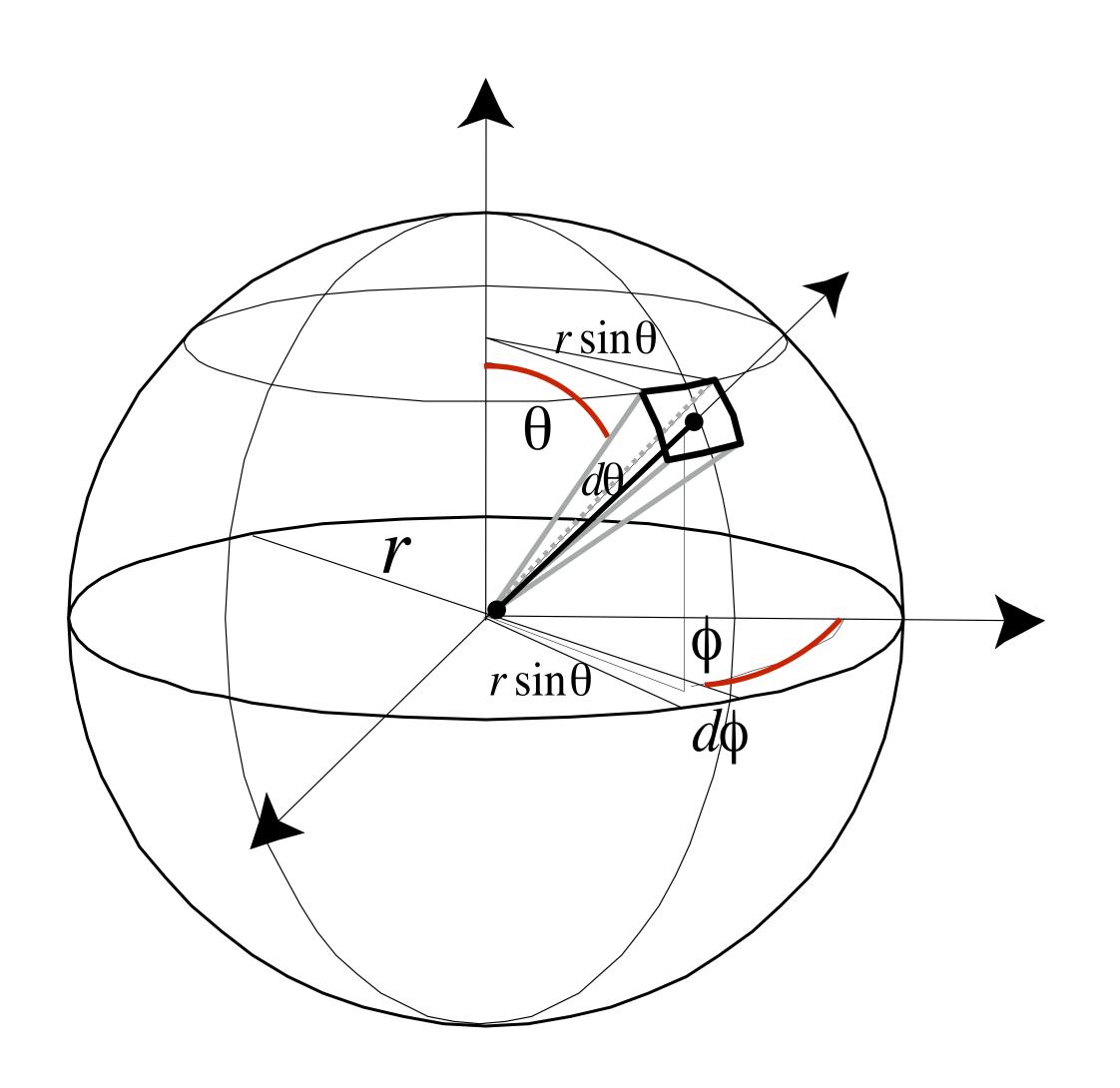






http://xkcd.com/1276/

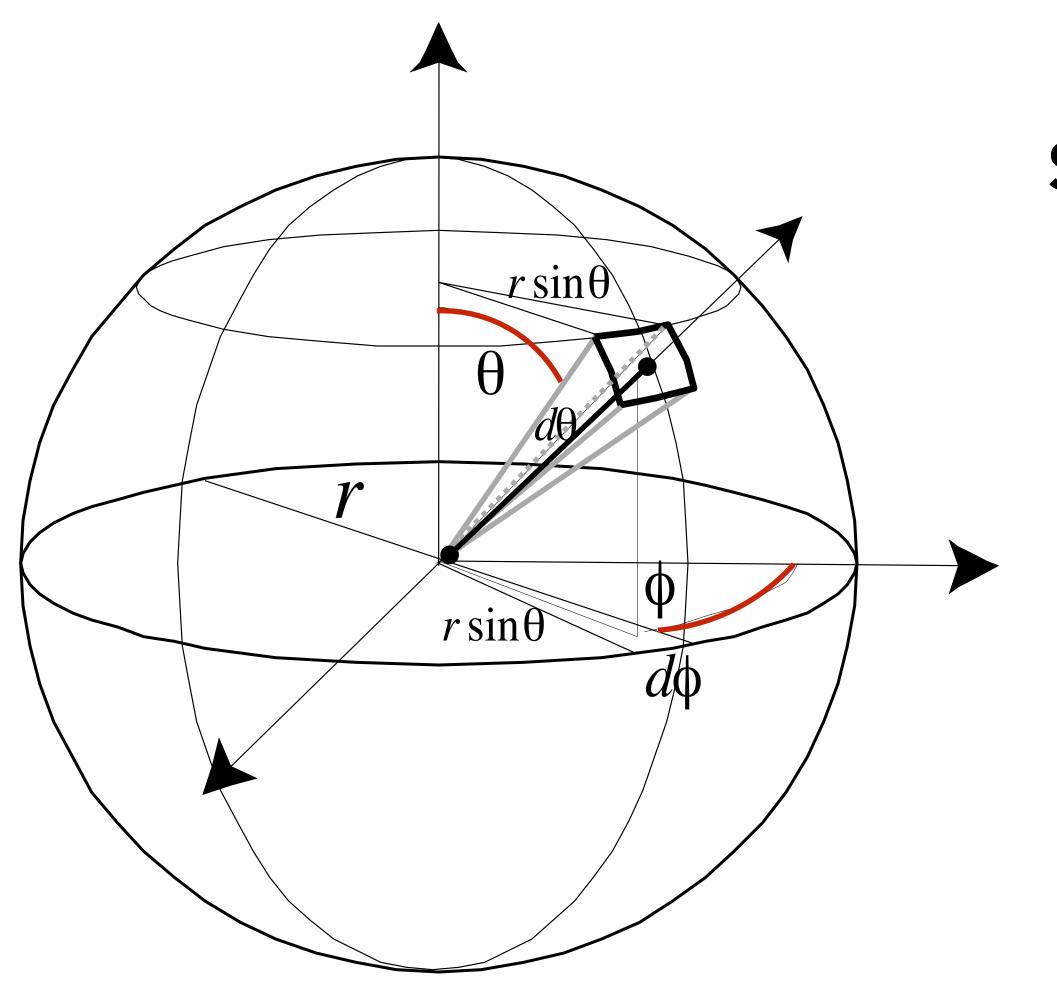
## Differential Solid Angles



$$dA = (r d\theta)(r \sin \theta d\phi)$$
$$= r^{2} \sin \theta d\theta d\phi$$

$$d\omega = \frac{dA}{r^2} = \sin\theta \, d\theta \, d\phi$$

## Differential Solid Angles



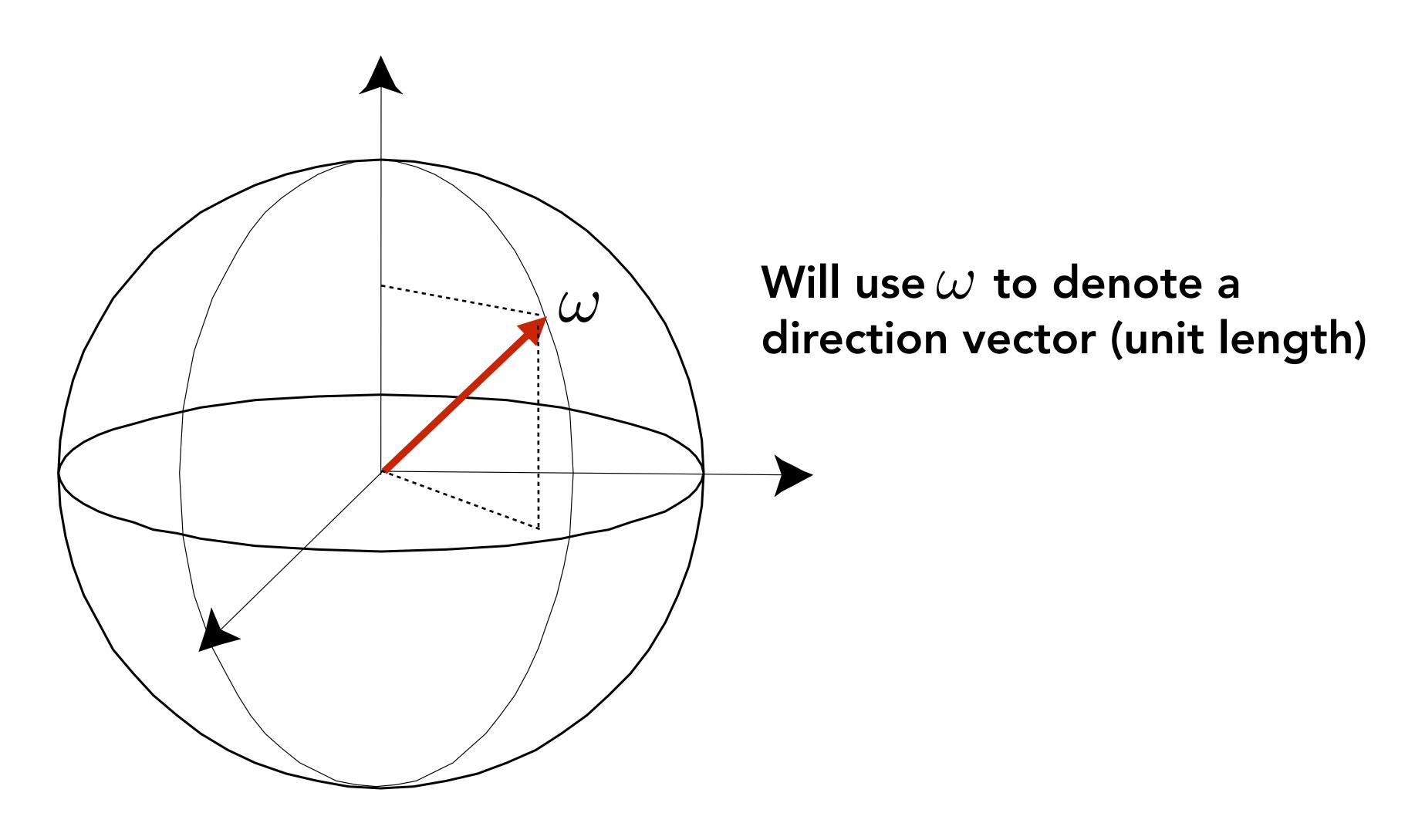
Sphere:  $S^2$ 

$$\Omega = \int_{S^2} d\omega$$

$$= \int_0^{2\pi} \int_0^{\pi} \sin \theta \, d\theta \, d\phi$$

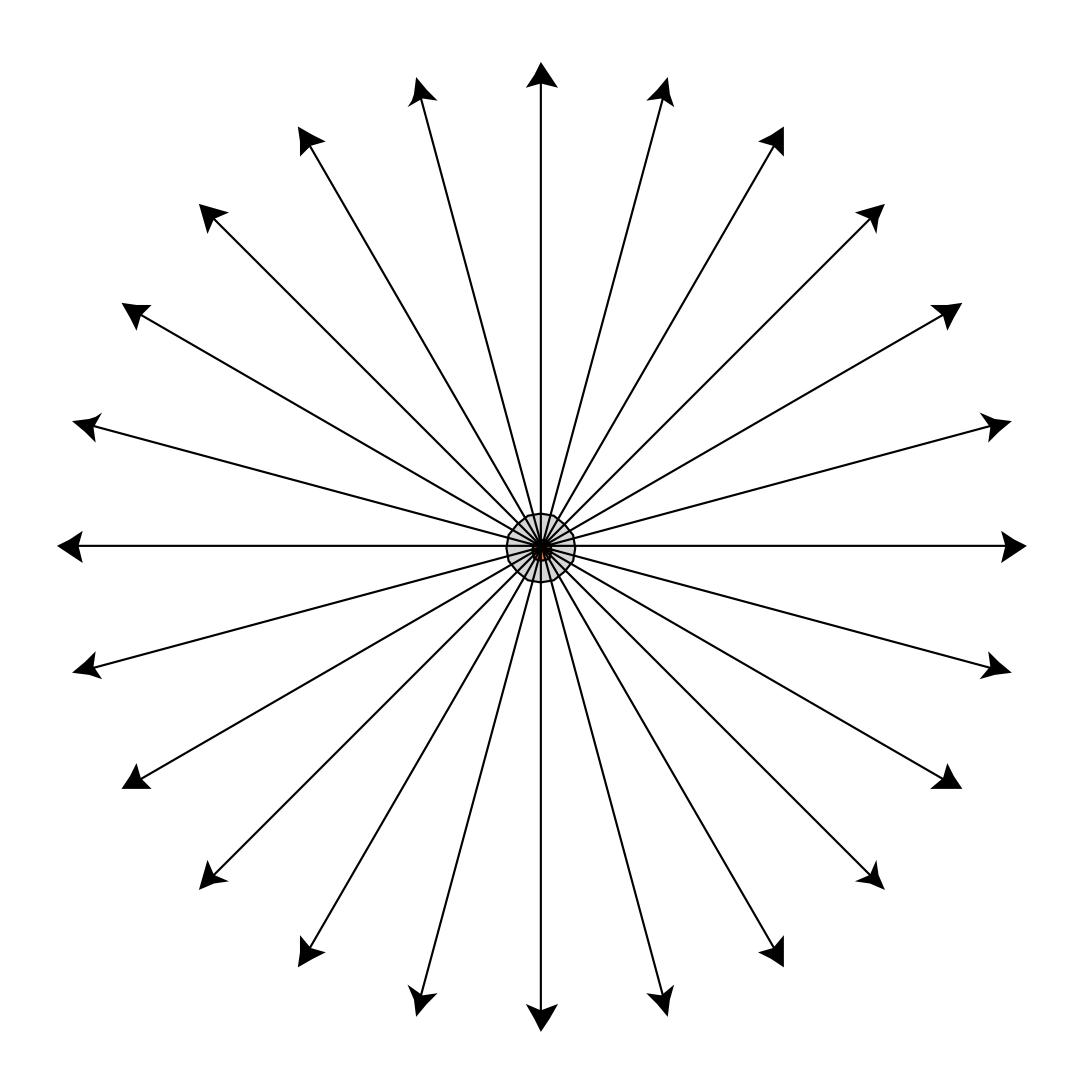
$$= 4\pi$$

#### $\omega$ as a direction vector



CS184/284A

## Isotropic Point Source



$$\Phi = \int_{S^2} I \, \mathrm{d}\omega$$
$$= 4\pi I$$

$$I = \frac{\Phi}{4\pi}$$

## Modern LED Light

Output: 815 lumens

(11W LED replacement for 60W incandescent)

Luminous intensity?

Assume isotropic: Intensity = 815 lumens / 4pi sr = 65 candelas

If focused into  $20^{\circ}$  diameter cone. Intensity = ??



## Spectral Power Distribution

#### Describes distribution of energy by wavelength

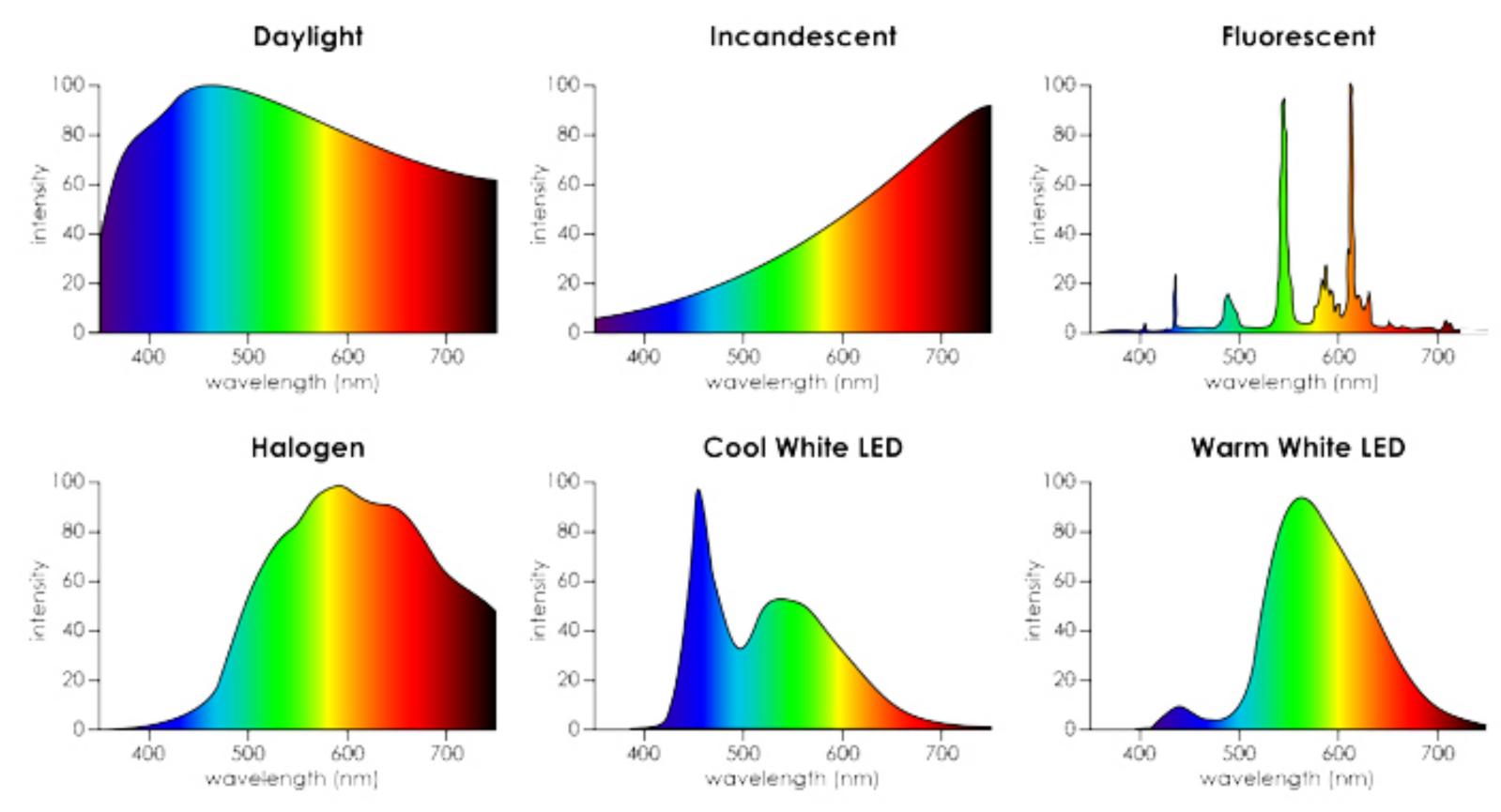
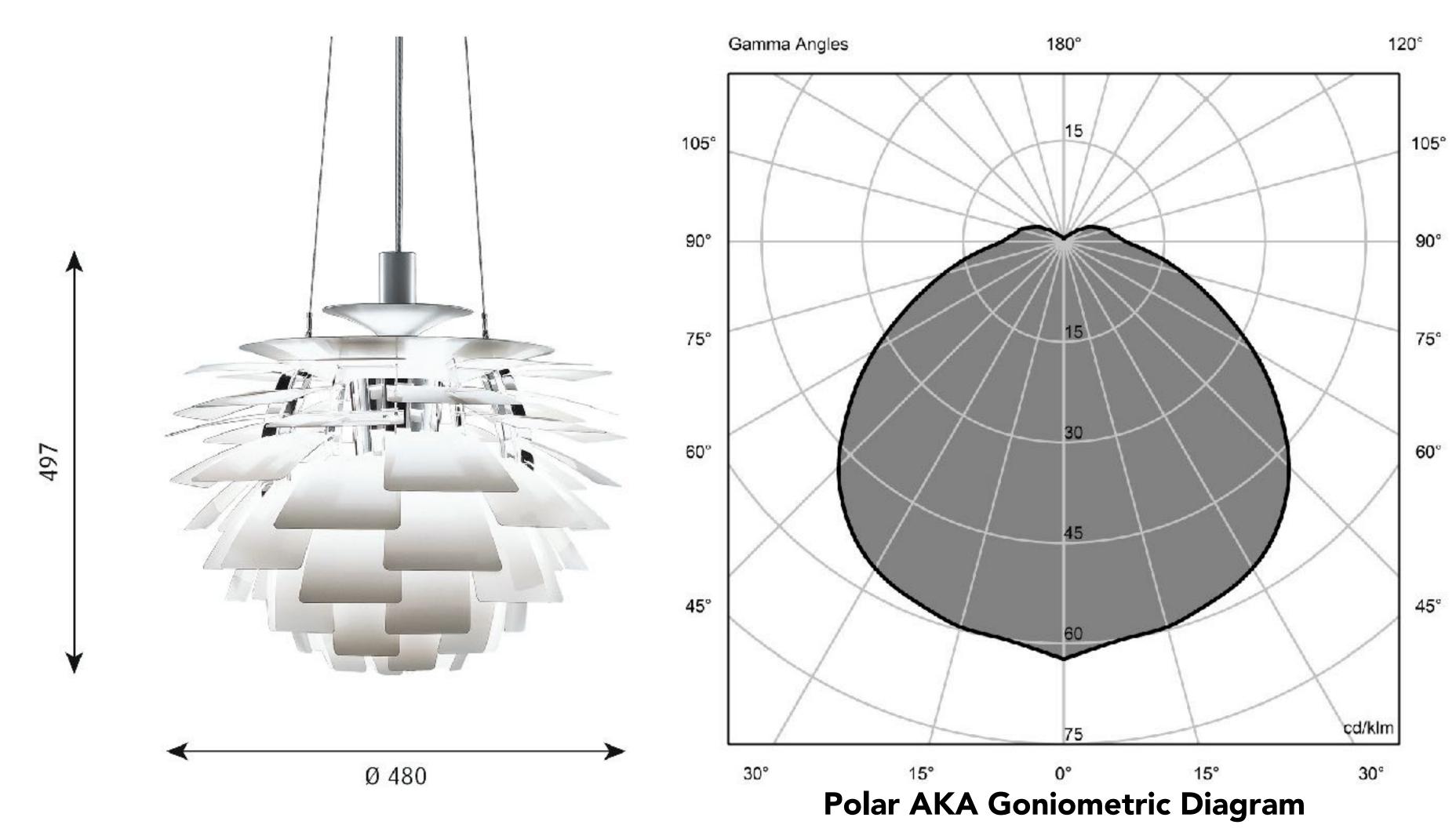


Figure credit:



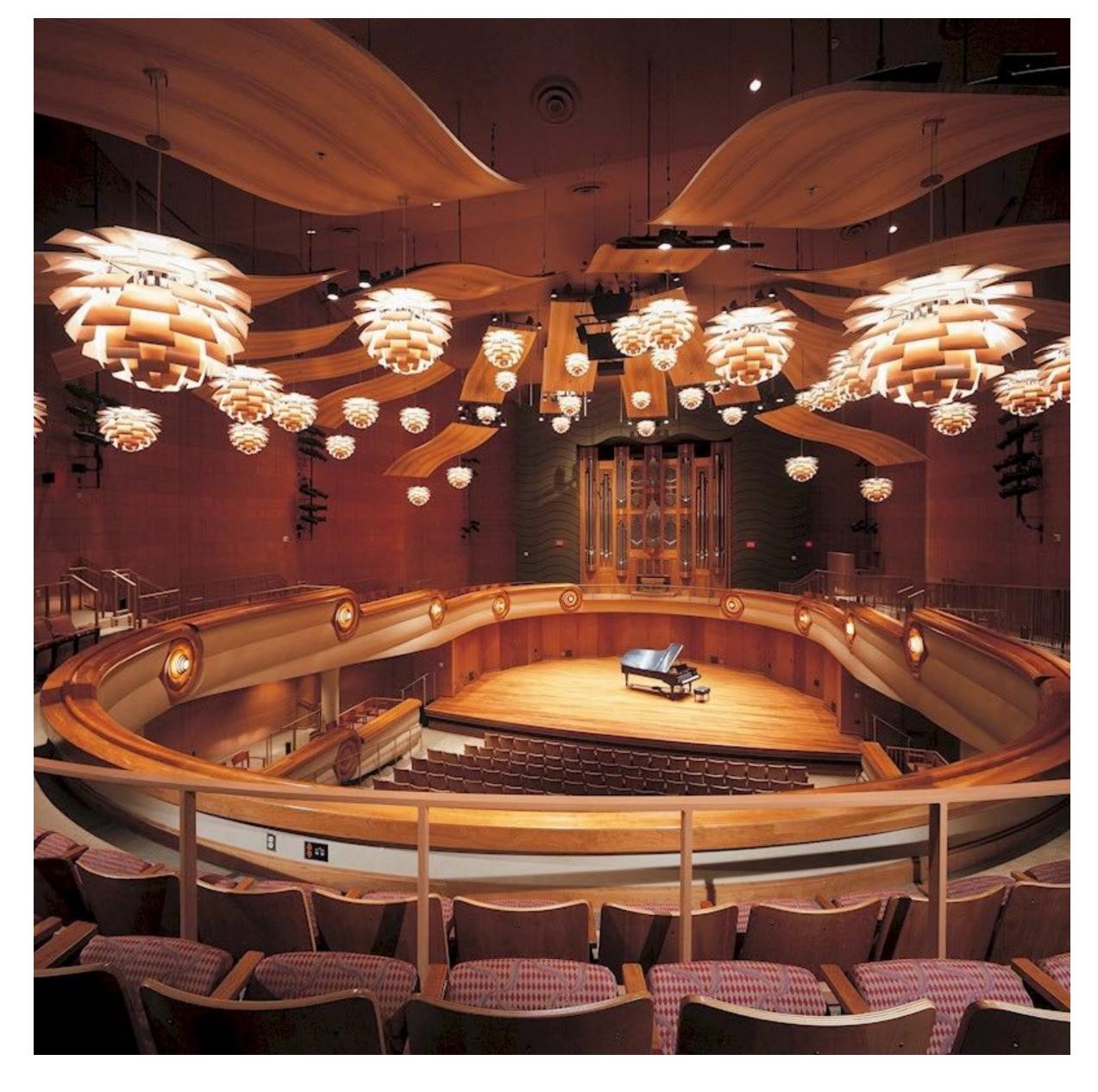
#### Light Fixture Measurements - Goniometric Diagram



Poul Henningsen's Artichoke Lamp

http://www.louispoulsen.com/

CS184/284A Ren Ng



PH Artichoke Lamps in Rivercenter for the Performing Arts, Georgia

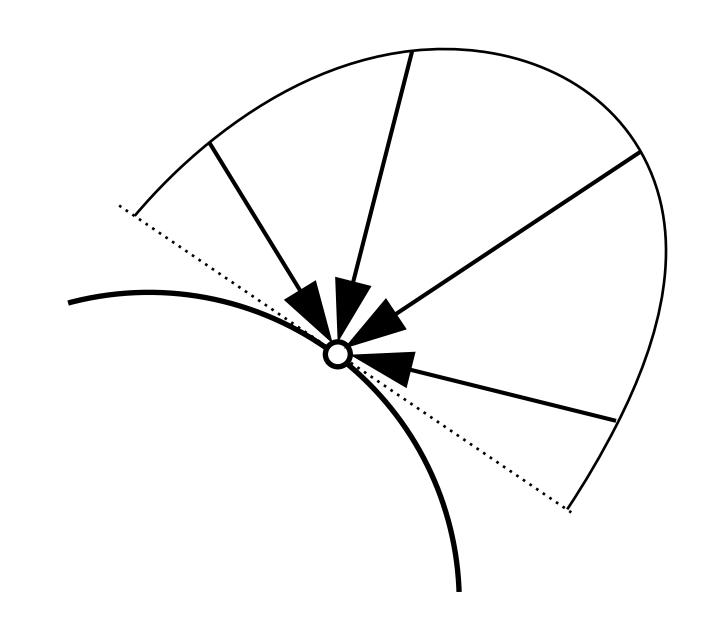


# Irradiance

#### Irradiance

Definition: The irradiance (illuminance) is the power per unit area incident on a surface point.

$$E(\mathbf{x}) \equiv \frac{\mathrm{d}\Phi(\mathbf{x})}{\mathrm{d}A}$$

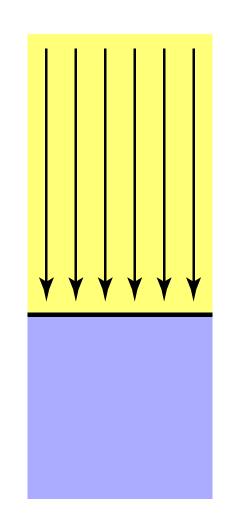


## Typical Values of Illuminance [lm/m<sup>2</sup>]

Brightest sunlight	120,000	lux
Overcast day (midday)	15,000	
Interior near window (daylight)	1,000	STATE OF THE PARTY
Residential artificial lighting	300	725 ' <b>95</b>
Sunrise / sunset	40	MONTH AND
Illuminated city street	10	MOLTONE MAGLEMAN
Moonlight (full)	0.02	Secrital Attacks
Starlight	0.0003	Light meter

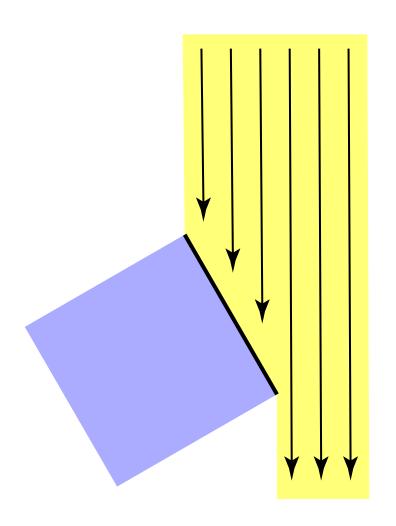
CS184/284A

#### Lambert's Cosine Law



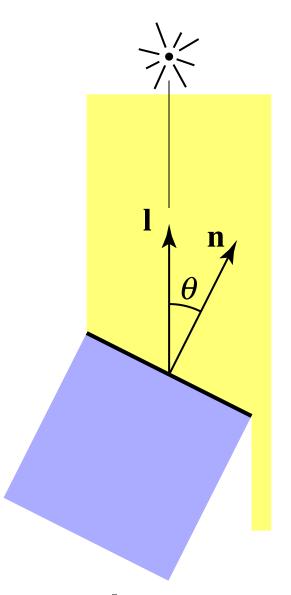
Top face of cube receives a certain amount of power

$$E = \frac{\Phi}{A}$$



Top face of 60° rotated cube receives half power

$$E = \frac{1}{2} \frac{\Phi}{A}$$



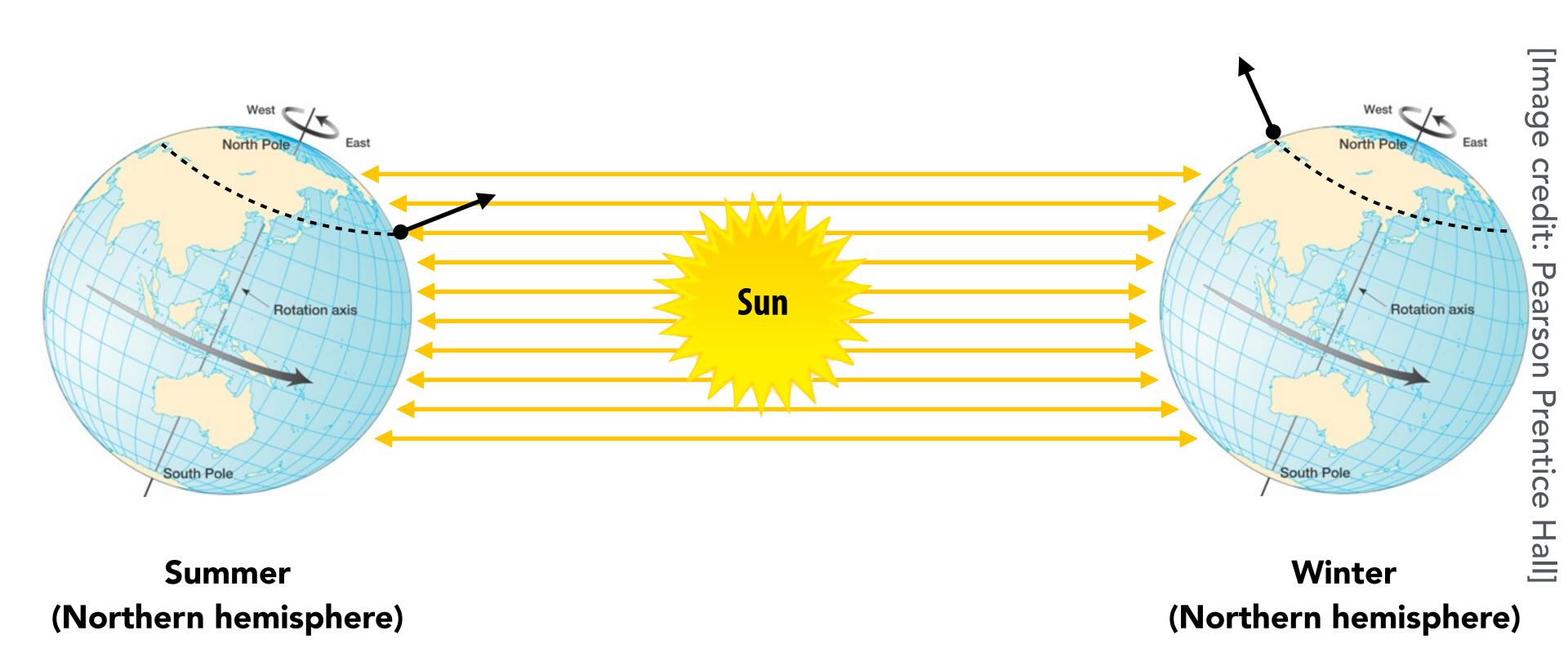
In general, power per unit area is proportional to

$$\cos\theta = l \cdot n$$

$$E = \frac{\Phi}{A}\cos\theta$$

Irradiance at surface is proportional to cosine of angle between light direction and surface normal.

## Why Do We Have Seasons?



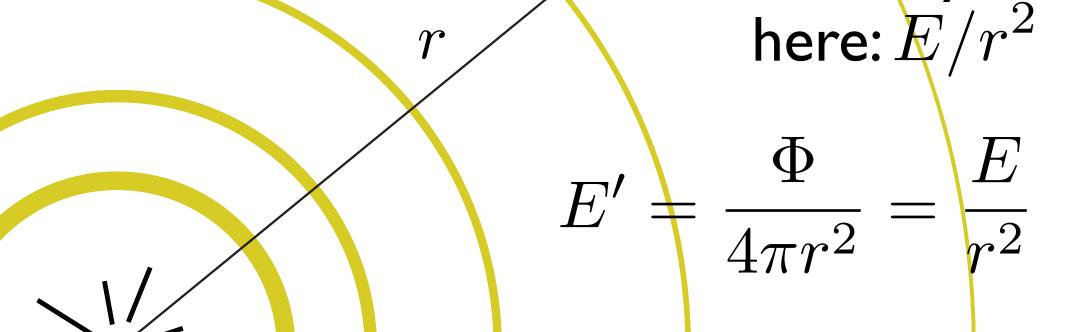
Earth's axis of rotation: ~23.5° off axis

CS184/284A Ren Ng

#### Irradiance Falloff

Assume light is emitting flux  $\Phi$  in a uniform angular distribution

Compare irradiance at surface of two spheres:



intensity

here: E

$$E = \frac{\Phi}{4\pi}$$

Ren Ng

intensity

#### Attendance

If you are seated in class, go to this form and sign in:

https://tinyurl.com/184lecture

#### Notes:

- Time-stamp will be taken when you submit form.
   Do it now, won't count later.
- Don't tell friends outside class to fill it out now, because we will audit at some point in semester.
- Failing audit will have large negative consequence.
   You don't need to, because you have an alternative!

# Radiance

#### Radiance

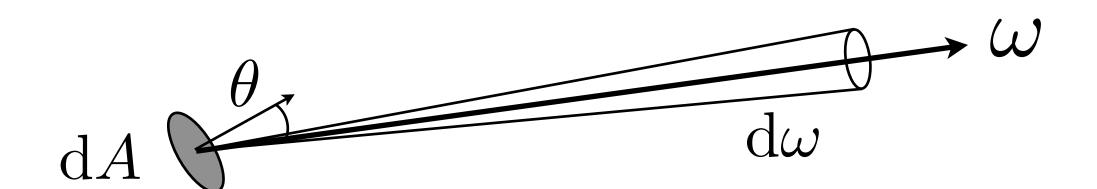


**Light Traveling Along A Ray** 

- 1. Radiance is the fundamental field quantity that describes the distribution of light in an environment
  - Radiance is the quantity associated with a ray
  - Rendering is all about computing radiance
- 2. Radiance is invariant along a ray in a vacuum

#### Surface Radiance

Definition: The radiance (luminance) is the power emitted, reflected, transmitted or received by a surface, per unit solid angle, per unit projected area.



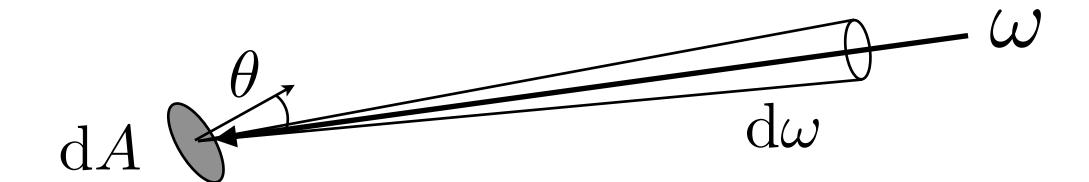
$$L(\mathbf{p}, \omega) \equiv \frac{\mathrm{d}^2 \Phi(\mathbf{p}, \omega)}{\mathrm{d}\omega \, \mathrm{d}A \cos \theta}$$

 $\cos \theta$  accounts for projected surface area

$$\left[\frac{W}{\operatorname{sr} m^2}\right] \left[\frac{\operatorname{cd}}{m^2} = \frac{\operatorname{lm}}{\operatorname{sr} m^2} = \operatorname{nit}\right]$$

#### Incident Surface Radiance

Equivalent: Incident surface radiance (luminance) is the irradiance per unit solid angle arriving at the surface.

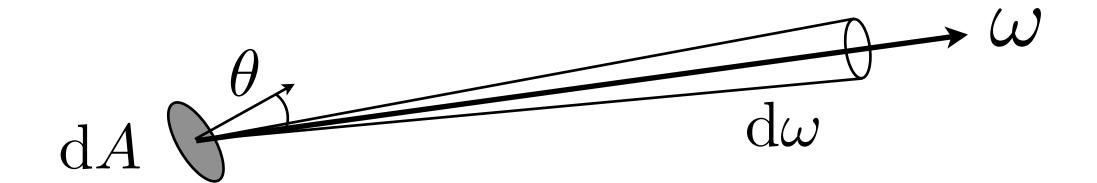


$$L(\mathbf{p}, \omega) = \frac{\mathrm{d}E(\mathbf{p})}{\mathrm{d}\omega \cos \theta}$$

i.e. it is the light arriving at the surface along a given ray (point on surface and incident direction).

## Exiting Surface Radiance

Equivalent: Exiting surface radiance (luminance) is the intensity per unit projected area leaving the surface.

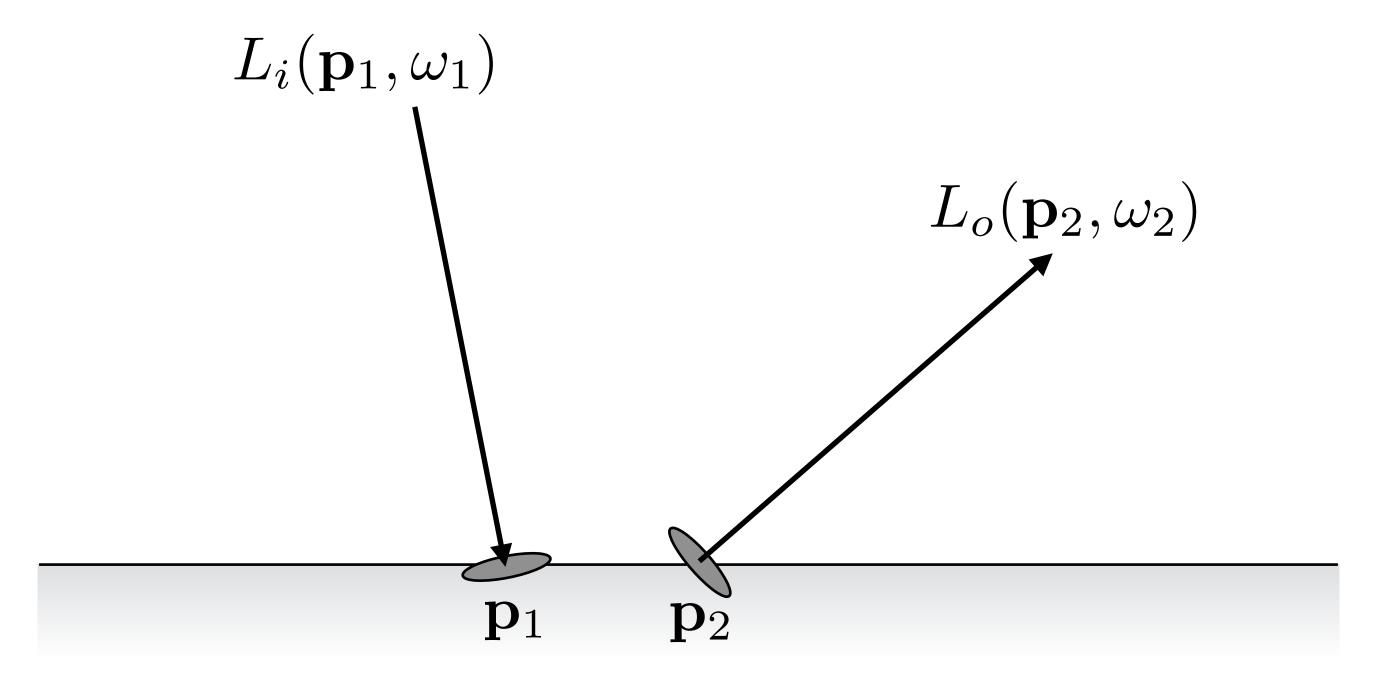


$$L(\mathbf{p}, \omega) = \frac{\mathrm{d}I(\mathbf{p}, \omega)}{\mathrm{d}A\cos\theta}$$

e.g. for an area light it is the light emitted along a given ray (point on surface and exit direction).

#### Incident & Exiting Surface Radiance Differ!

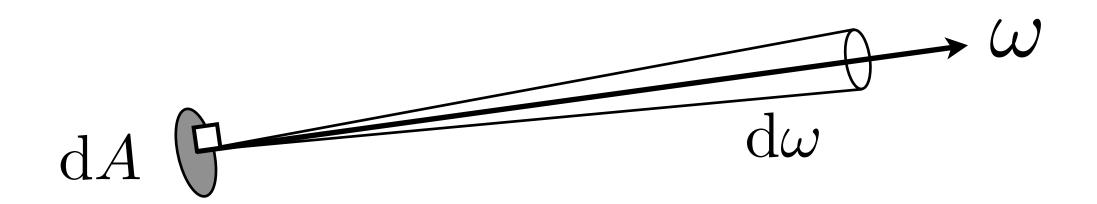
Need to distinguish between incident radiance and exitant radiance functions at a point on a surface



In general:  $L_i(\mathbf{p},\omega) \neq L_o(\mathbf{p},\omega)$ 

### Field Radiance or Light Field

Definition: The field radiance (luminance) at a point in space in a given direction is the power per unit solid angle per unit area perpendicular to the direction.



CS184/284A

### Typical Values of Luminance [cd/m²]

Surface of the sun	2,000,000,000 nits	>
Sunlight clouds	30,000	
Clear sky	3,000	
Cellphone display	500	
Overcast sky	300	
Scene at sunrise	30	
Scene lit by moon	0.001	
Threshold of vision	0.00001	

CS184/284A

# Calculating with Radiance

#### Irradiance from the Environment

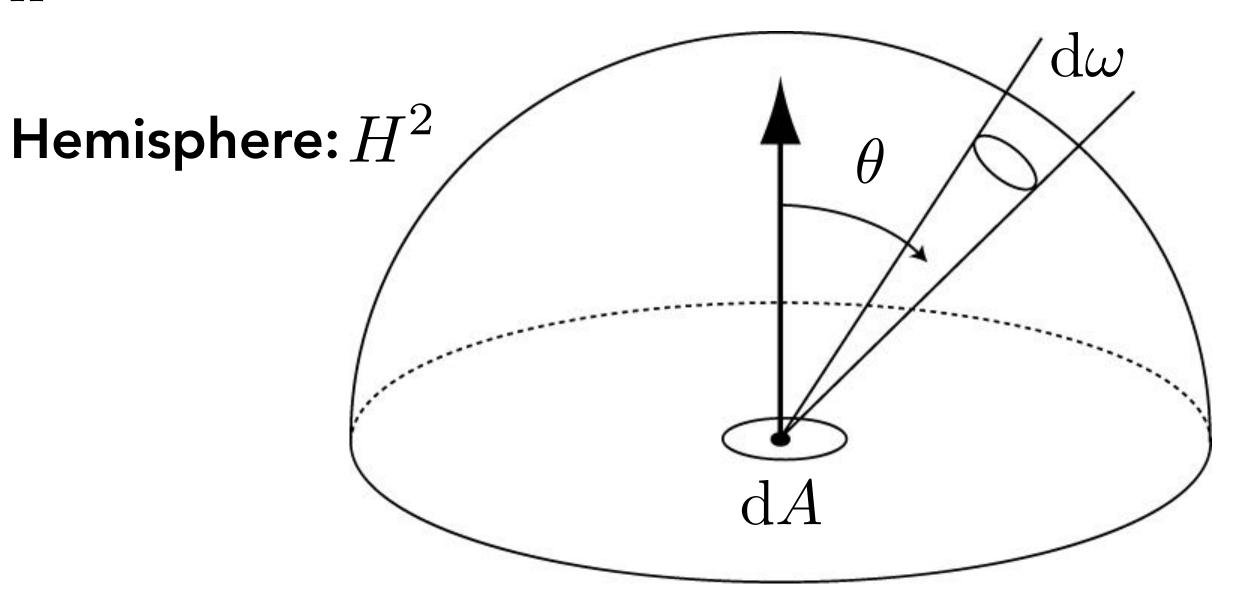
Computing flux per unit area on surface, due to incoming light from all directions.

$$dE(\mathbf{p},\omega) = L_i(\mathbf{p},\omega)\cos\theta\,\mathrm{d}\omega \qquad \qquad \text{Contribution to irradiance from light arriving from direction }\omega$$

$$E(\mathbf{p}) = \int_{H^2} L_i(\mathbf{p}, \omega) \cos \theta \, d\omega$$



Light meter



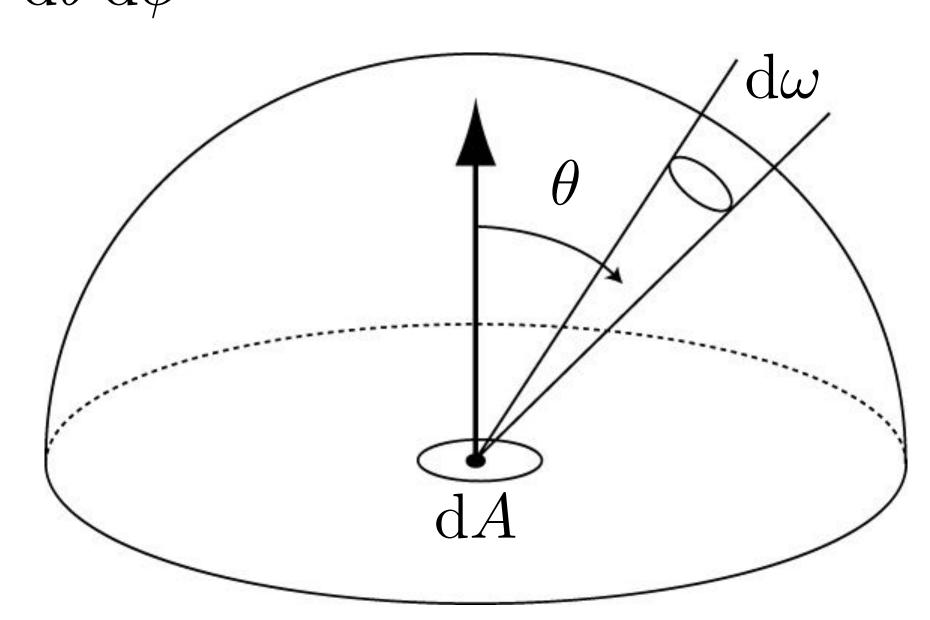
#### Irradiance from Uniform Hemispherical Light

$$E(\mathbf{p}) = \int_{H^2} L \cos \theta \, d\omega$$

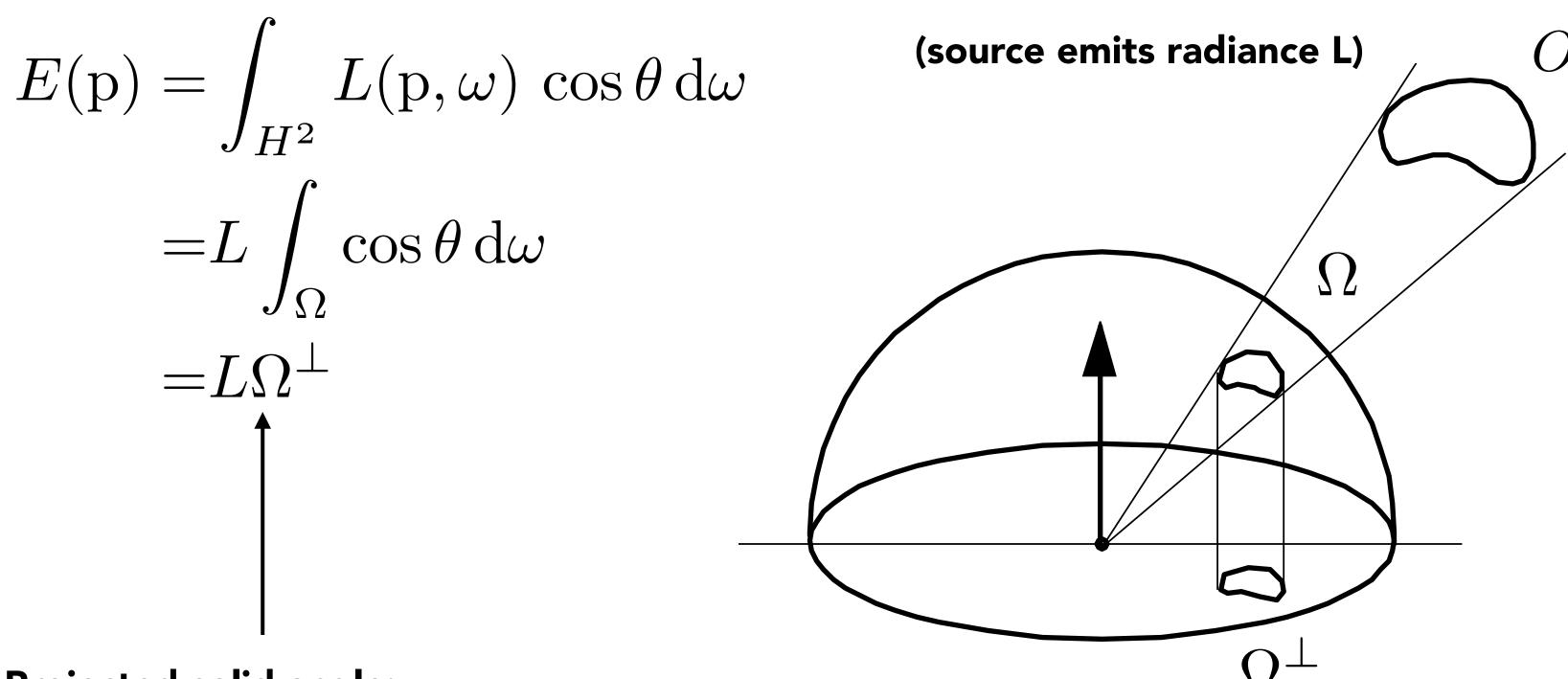
$$= L \int_0^{2\pi} \int_0^{\frac{\pi}{2}} \cos \theta \sin \theta \, d\theta \, d\phi$$

$$= L \pi$$

Note: integral of cosine over hemisphere is only 1/2 the area of the hemisphere.



#### Irradiance from a Uniform Area Source



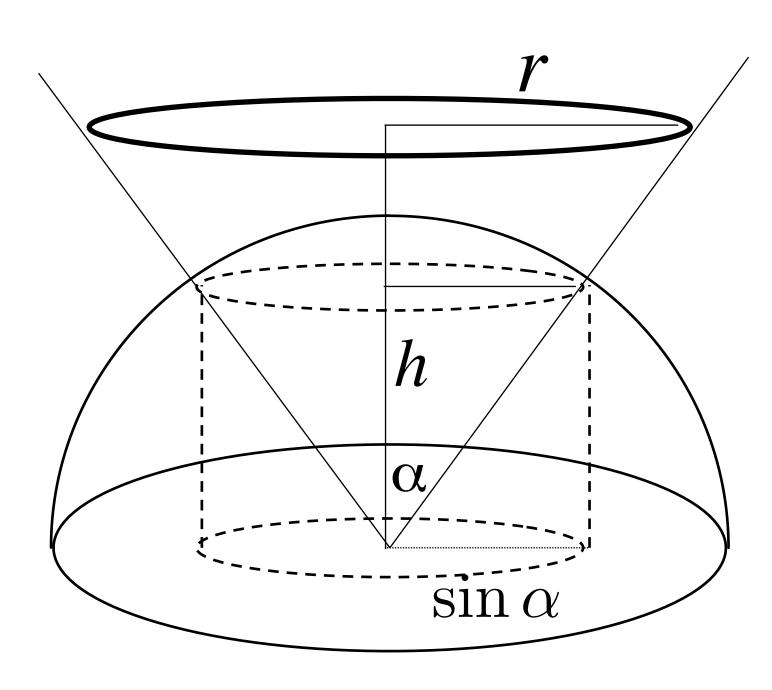
#### Projected solid angle:

- Cosine-weighted solid angle
- Area of object O projected onto unit sphere, then projected onto plane

$$d\omega^{\perp} = |\cos\theta| d\omega$$

#### Uniform Disk Source Overhead

#### **Geometric Derivation**



$$\Omega^{\perp} = \pi \sin^2 \alpha$$

#### **Algebraic Derivation**

$$\Omega^{\perp} = \int_{0}^{2\pi} \int_{0}^{\alpha} \cos \theta \sin \theta \, d\theta \, d\phi$$

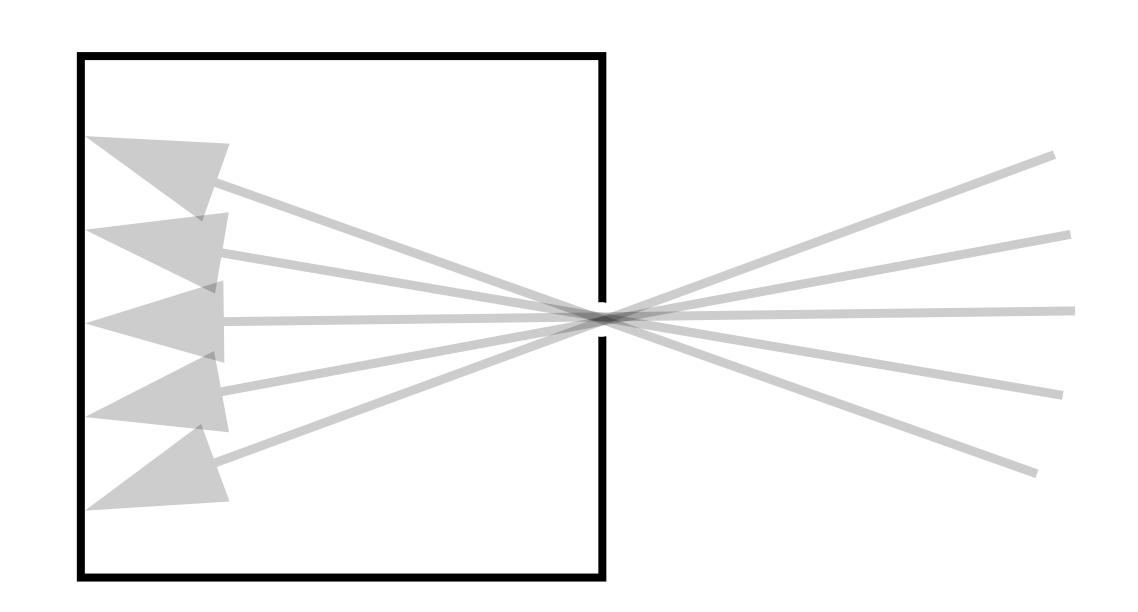
$$= 2\pi \frac{\sin^{2} \theta}{2} \Big|_{0}^{\alpha}$$

$$= \pi \sin^{2} \alpha$$

# Measuring Radiance

#### A Pinhole Camera Samples Radiance

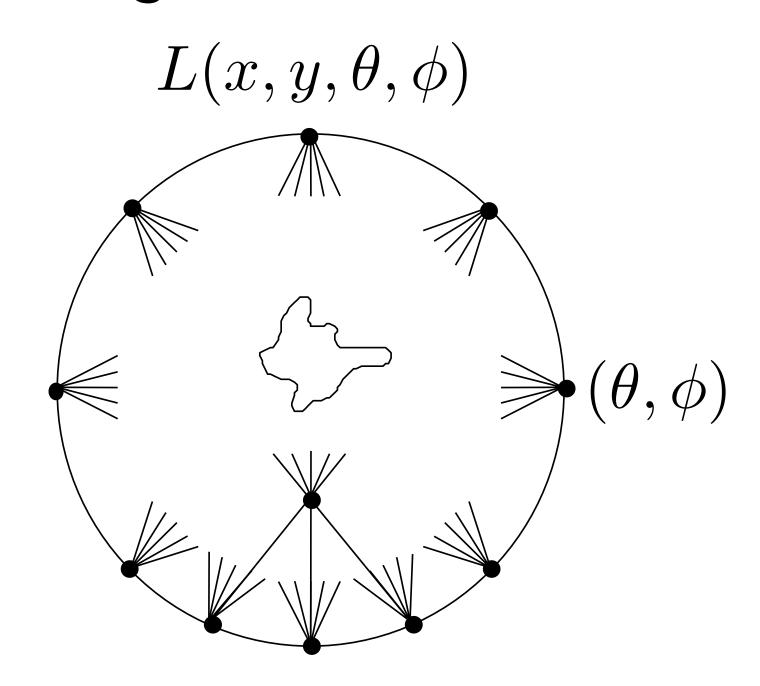
Photograph pixels measure radiance for rays passing through pinhole in different directions

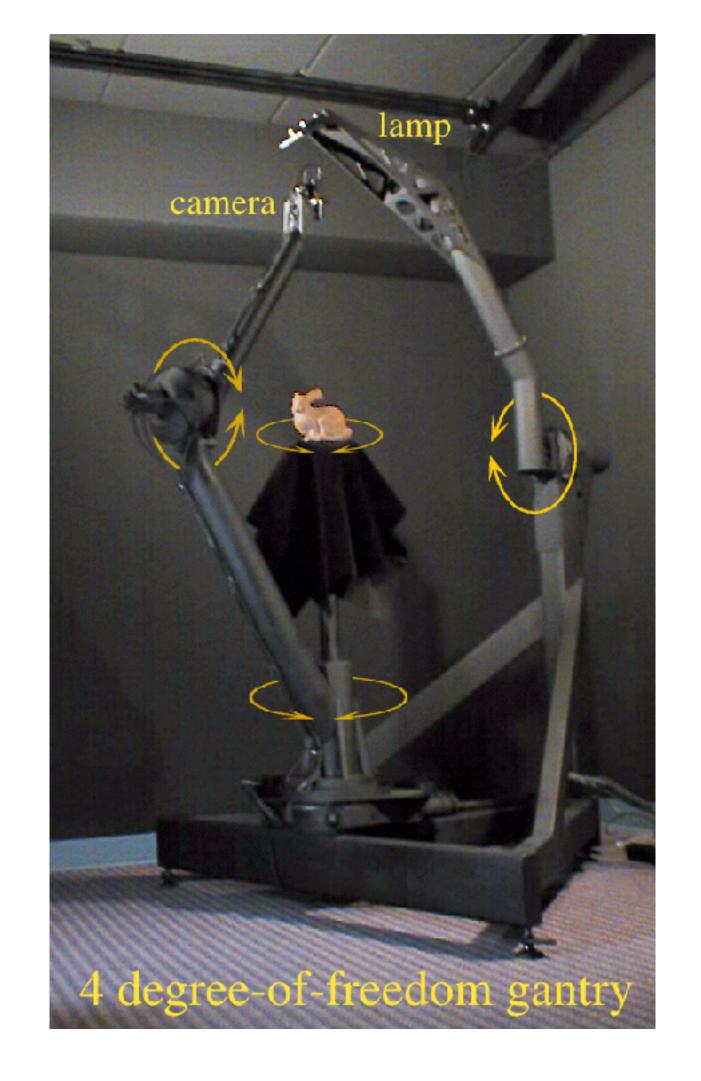


### Spherical Gantry ⇒ 4D Light Field

Take photographs of an object from all points on an enclosing sphere

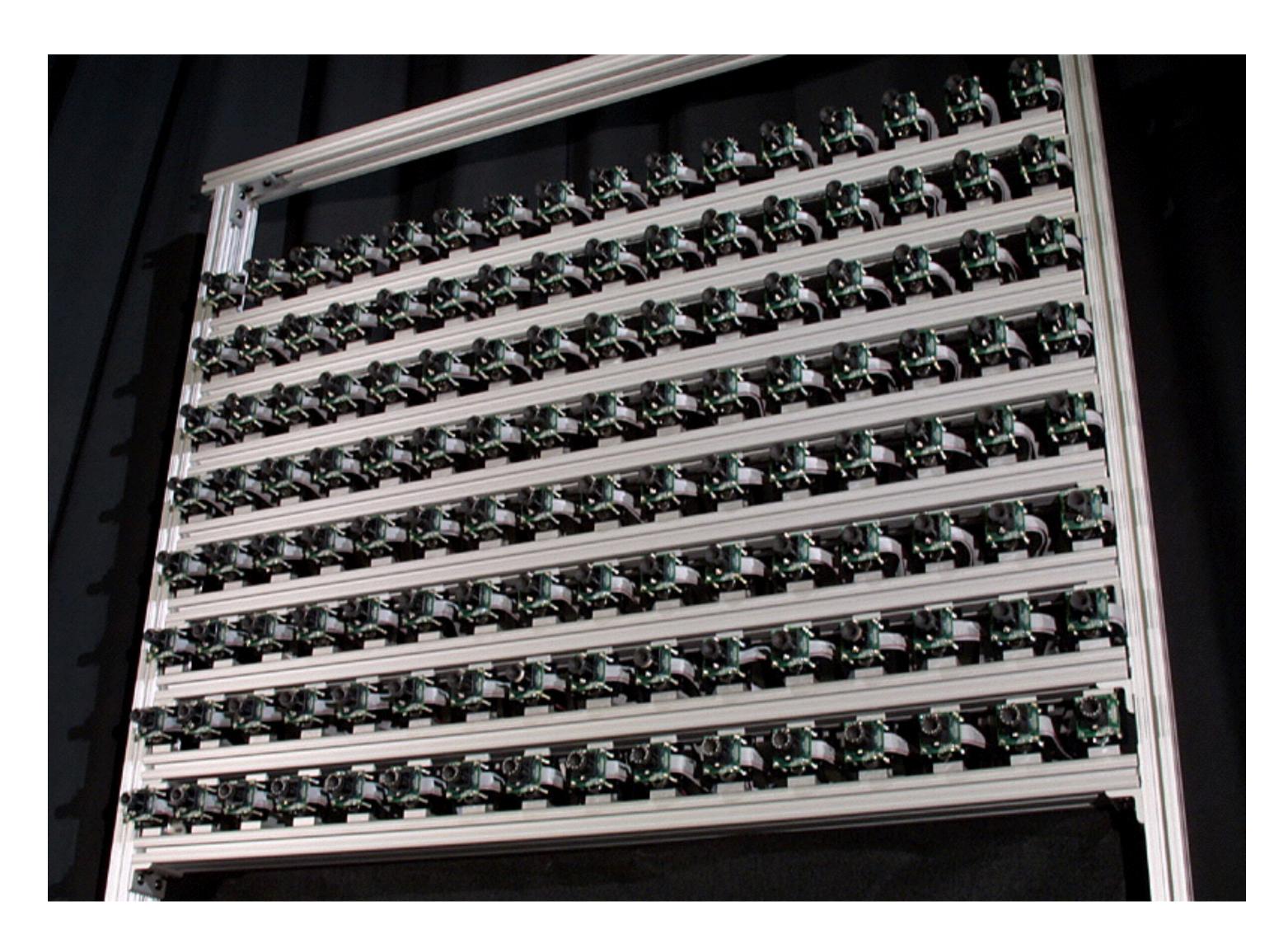
Captures all light leaving an object – like a hologram



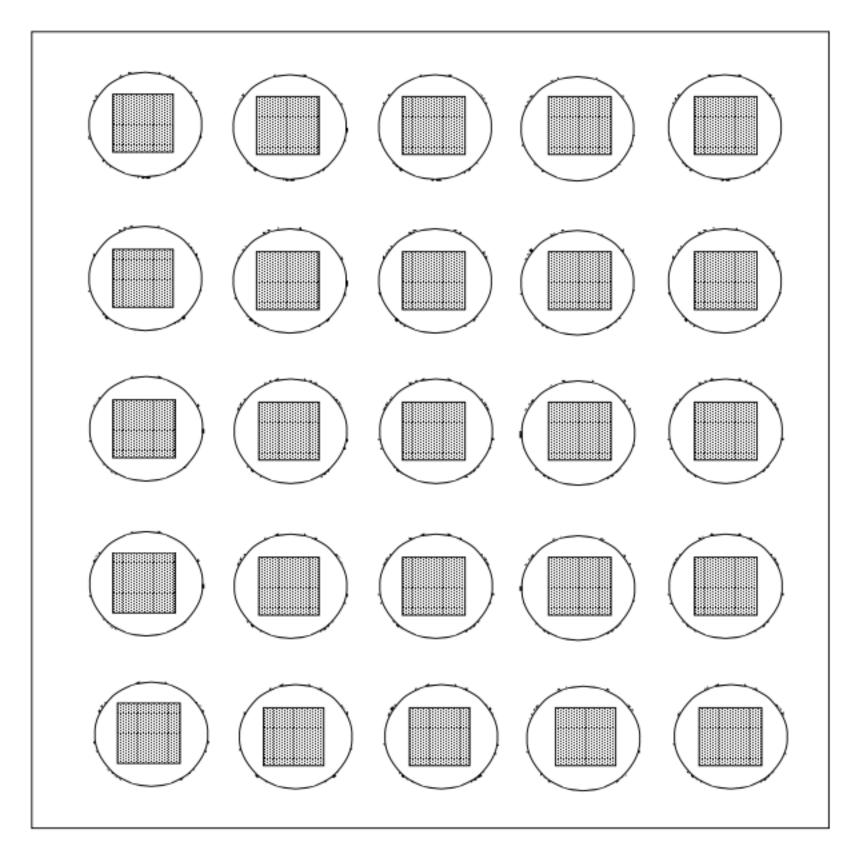


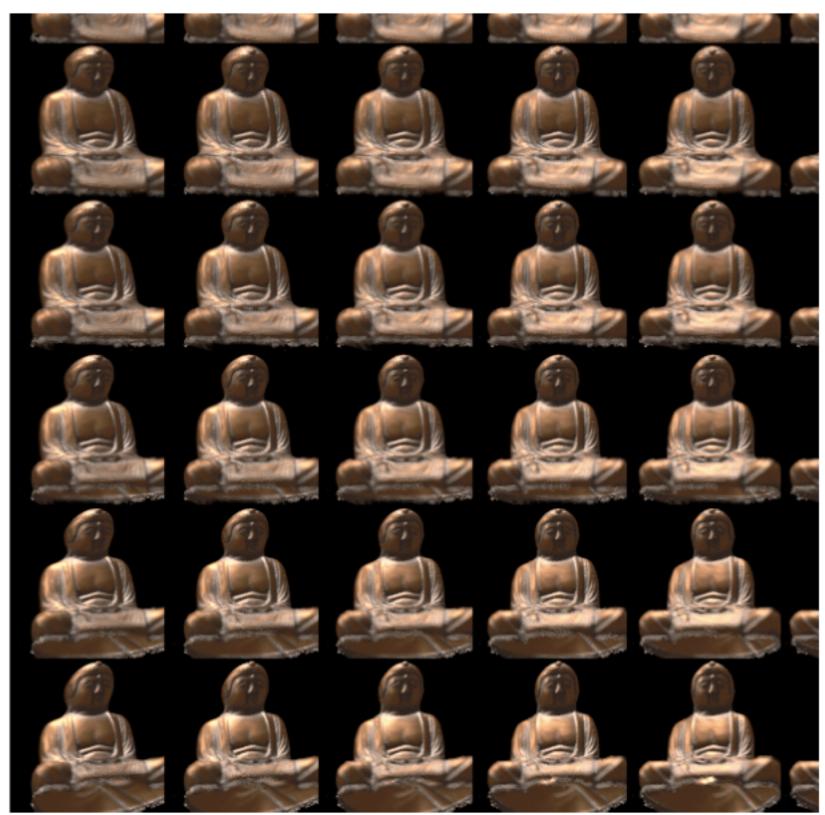
CS184/284A

# Multi-Camera Array ⇒ 4D Light Field



### Two-Plane Light Field





2D Array of Cameras

2D Array of Images

# Radiometry & Photometry Terms & Units

#### Radiometric & Photometric Terms & Units

Physics		Radiometry	Units	Photometry	Units
Energy	$\overline{Q}$	Radiant Energy	Joules (W·sec)	Luminous Energy	Lumen·sec
Flux (Power)	Φ	Radiant Power	W	Luminous Power	Lumen (Candela sr)
Angular Flux Density	I	Radiant Intensity	W/sr	Luminous Intensity	Candela (Lumen/sr)
Spatial Flux Density	E	Irradiance (in) Radiosity (out)	W/m <sup>2</sup>	Illuminance (in) Luminosity (out)	Lux (Lumen/m²)
Spatio-Angular Flux Density	L	Radiance	W/m <sup>2</sup> /sr	Luminance	Nit (Candela/m²)

<sup>&</sup>quot;Thus one nit is one lux per steradian is one candela per square meter is one lumen per square meter per steradian. Got it?" — James Kajiya

### Things to Remember

Radiometry vs photometry: physics vs human response Spatial measures of light:

- Flux, intensity, irradiance, radiance
- Pinhole cameras and light field cameras

Lighting calculations

- Integration on sphere / hemisphere
- Cosine weight: project from hemisphere onto disk
- Photon counting

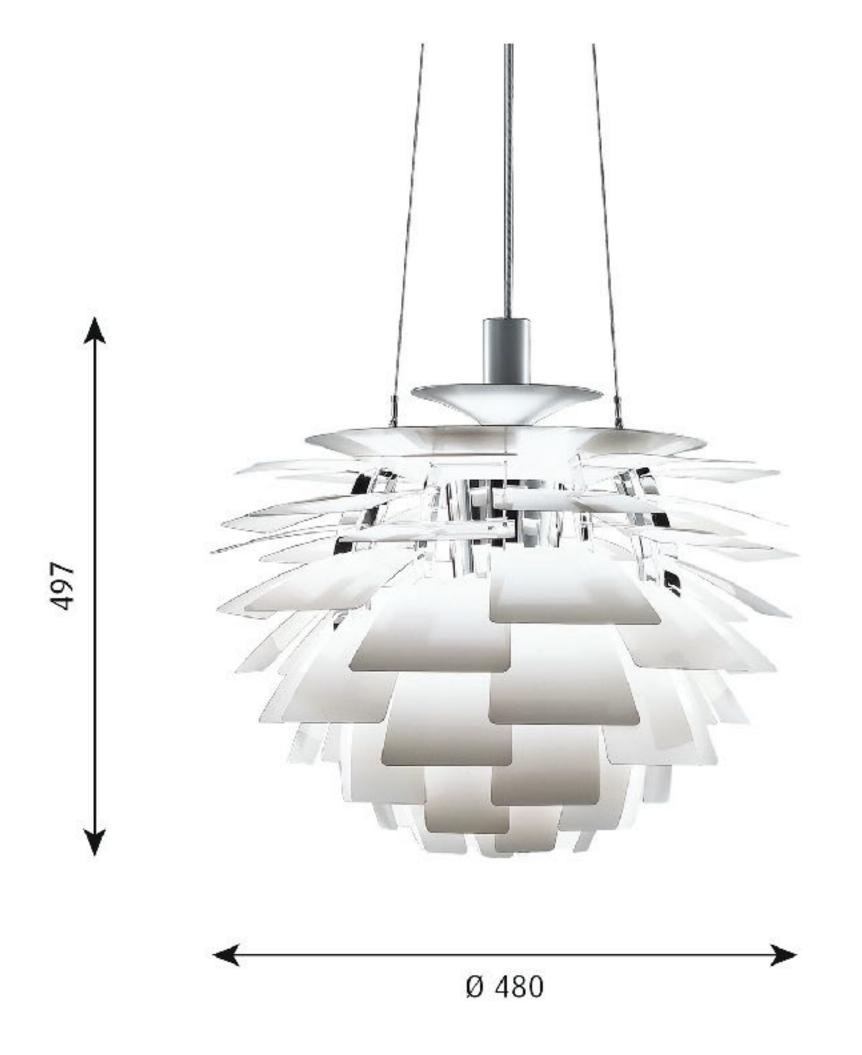
BRDF: 4D function for material reflection at a point

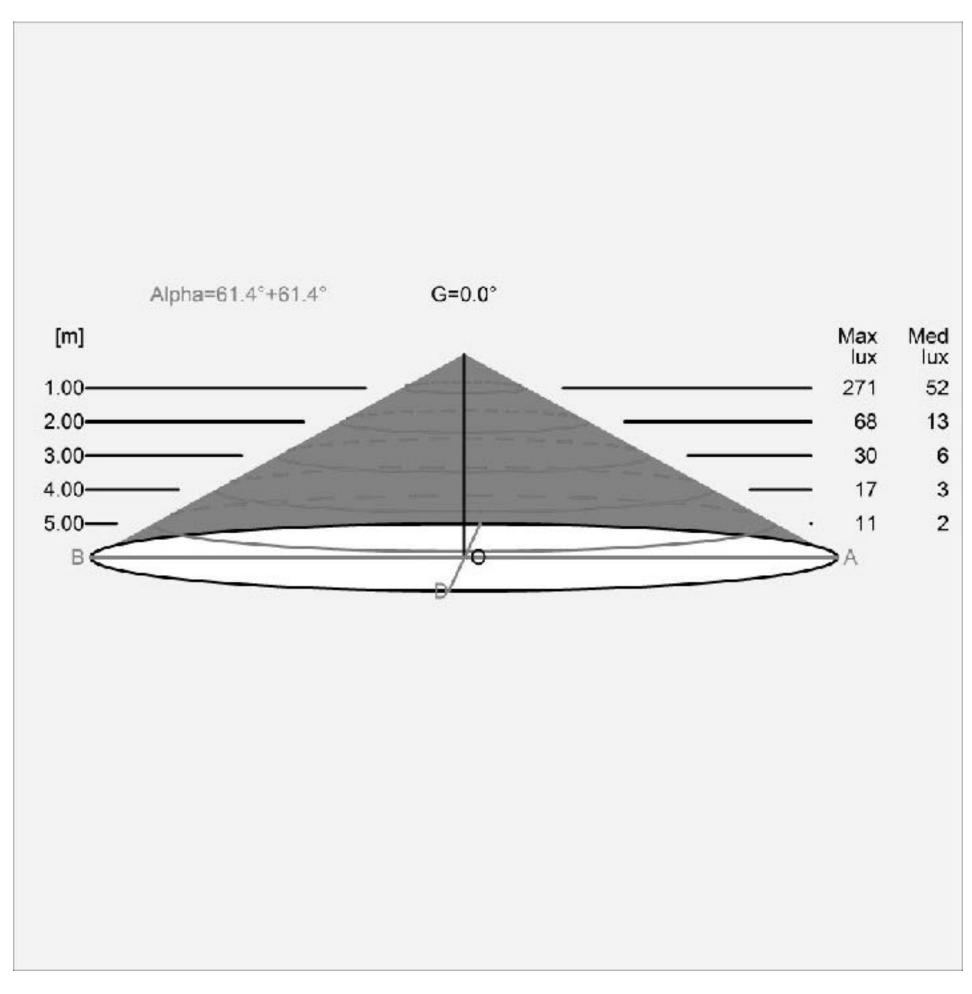
### Acknowledgments

Many thanks to Kayvon Fatahalian, Matt Pharr, Pat Hanrahan, and Steve Marschner for presentation resources.

### Extra

### Light Fixture Measurements



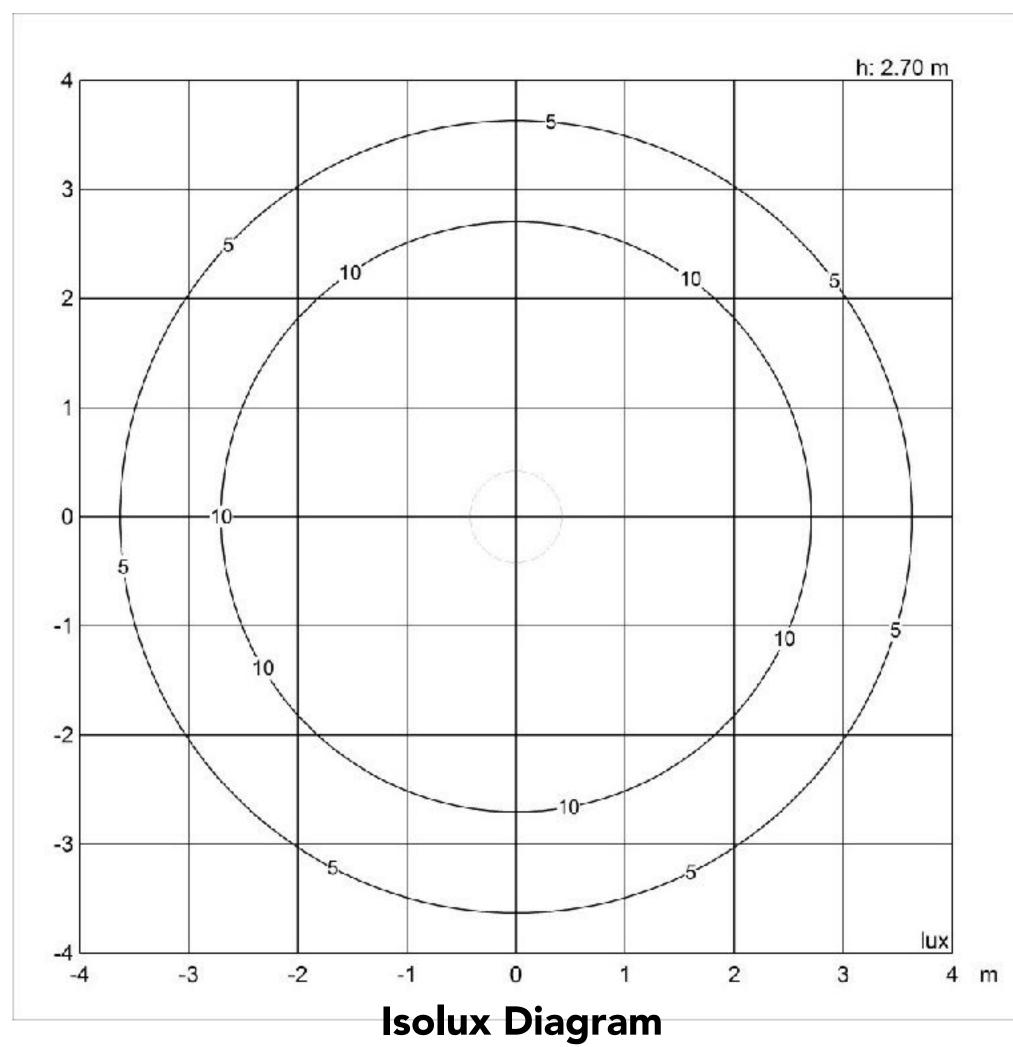


Poul Henningsen's Artichoke Lamp

Cartesian Diagram
http://www.louispoulsen.com/

# Light Fixture Measurements





Poul Henningsen's Artichoke Lamp

http://www.louispoulsen.com/

# Quantitative Photometry

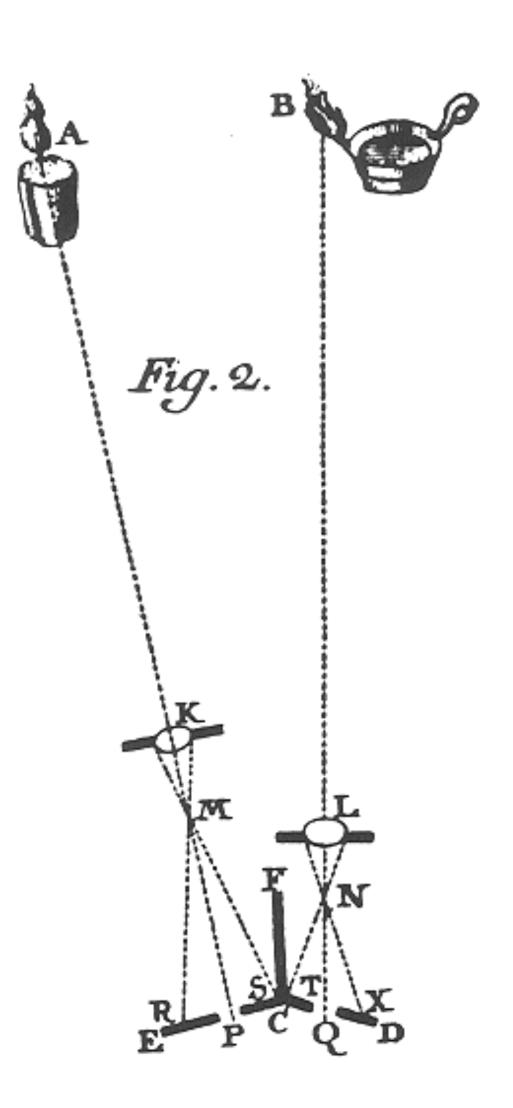
### The Invention of Photometry

#### Bouguer's classic experiment

- Compare a light source and a candle
- Move until appear equally bright
- Intensity is proportional to ratio of distances squared

#### Definition of a candela

- Originally a "standard" candle
- Currently 555 nm laser with power 1/683 W/sr
- One of seven SI base units



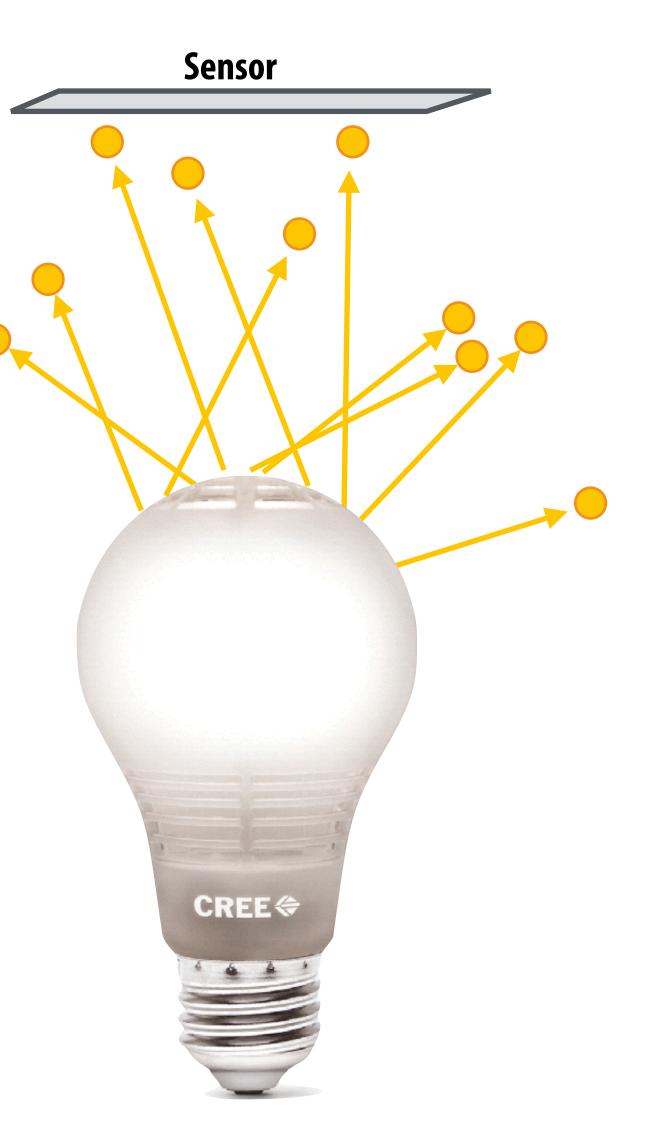
# Counting Photons

Given a sensor/light, we can count how many photons it receives/emits

- Over a period of time, gives the energy Q and flux (power)  $\Phi$  received/emitted by the sensor/light
- Energy carried by a photon:

$$Q=\frac{hc}{\lambda}$$
, where  $h\approx 6.626\times 10^{-34} \mathrm{m^2\,kg/s}$   $c=299,792,458~\mathrm{m/s}$   $\lambda=\mathrm{wavelength}$  of photon

- ~ 3.6 E-19J for a 555nm green photon
- ~ 2.8 E18 green photons for 1W of radiant energy



### Modern LED Light: Estimate Efficiency?

Input power: 11 W

Output: 815 lumens

(~80 lumens / Watt)

Incandescent bulb?

Input power: 60W

Output: ~700 lumens

(~12 lumens / Watt)



CS184/284A

### Modern LED Light: Estimate Efficiency?

Input power: 11 W
If all power into light with
555nm average wavelength,
get 3.1E19 photons/s

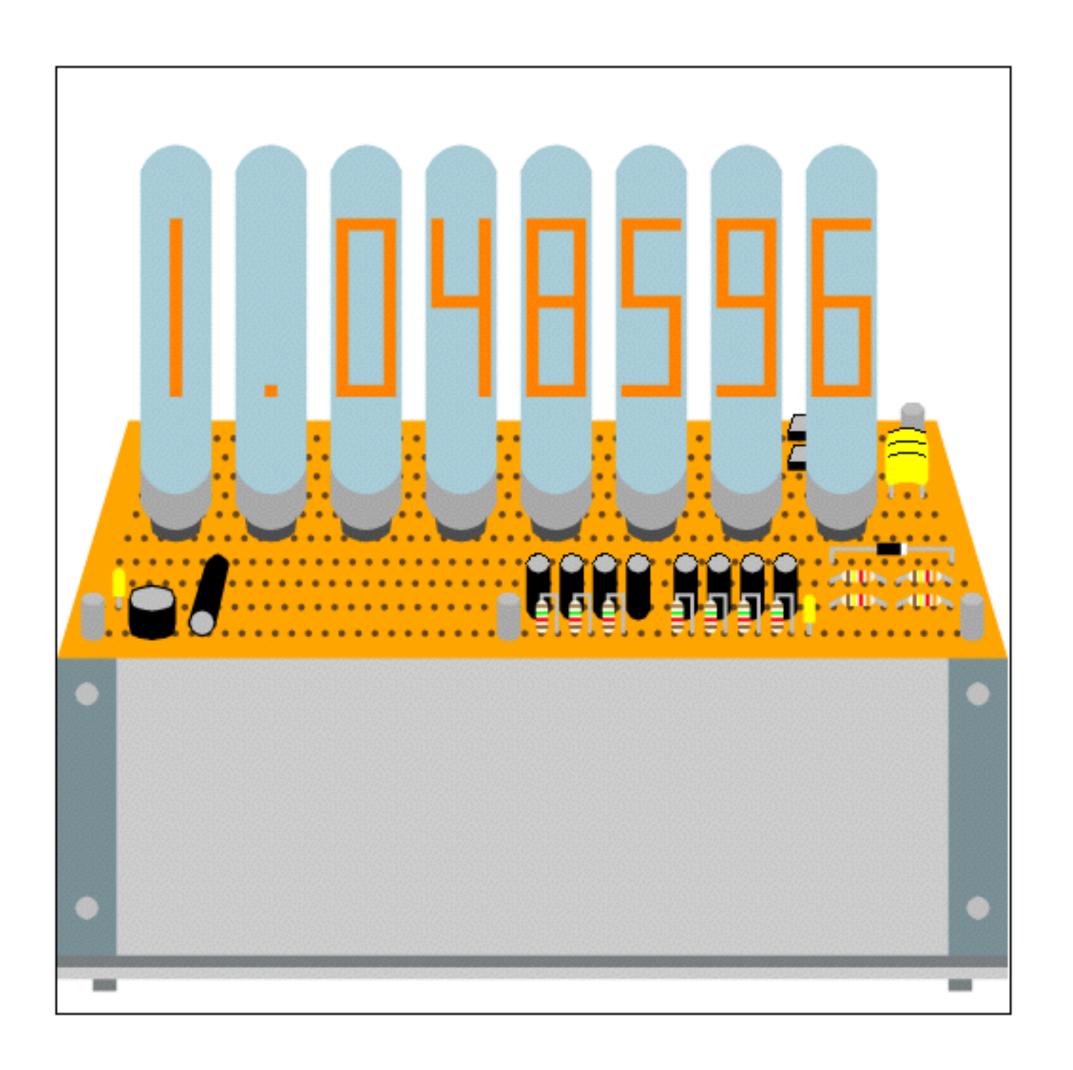
Intensity rating is 815 lumens, equivalent to 555nm laser at 815/683W. If average wavelength is 555nm, get 3.3E18 photons/s.

Efficiency\*: 3.3E18/3.1E19 = 11%



# Art Competition #1 Results

#### Art Competition #1 – 3rd Place Winner



#### Shannon Hu

It's a bunch of Nixie tubes sitting on a perforated circuit board with some other electronic components (resistors, capacitors, etc.).

#### Approach:

I manually wrote the SVG code for it, using a Google Image as reference for the dimensions. I used mostly polygons and polylines.

#### Art Competition #1 – 2nd Place Winner



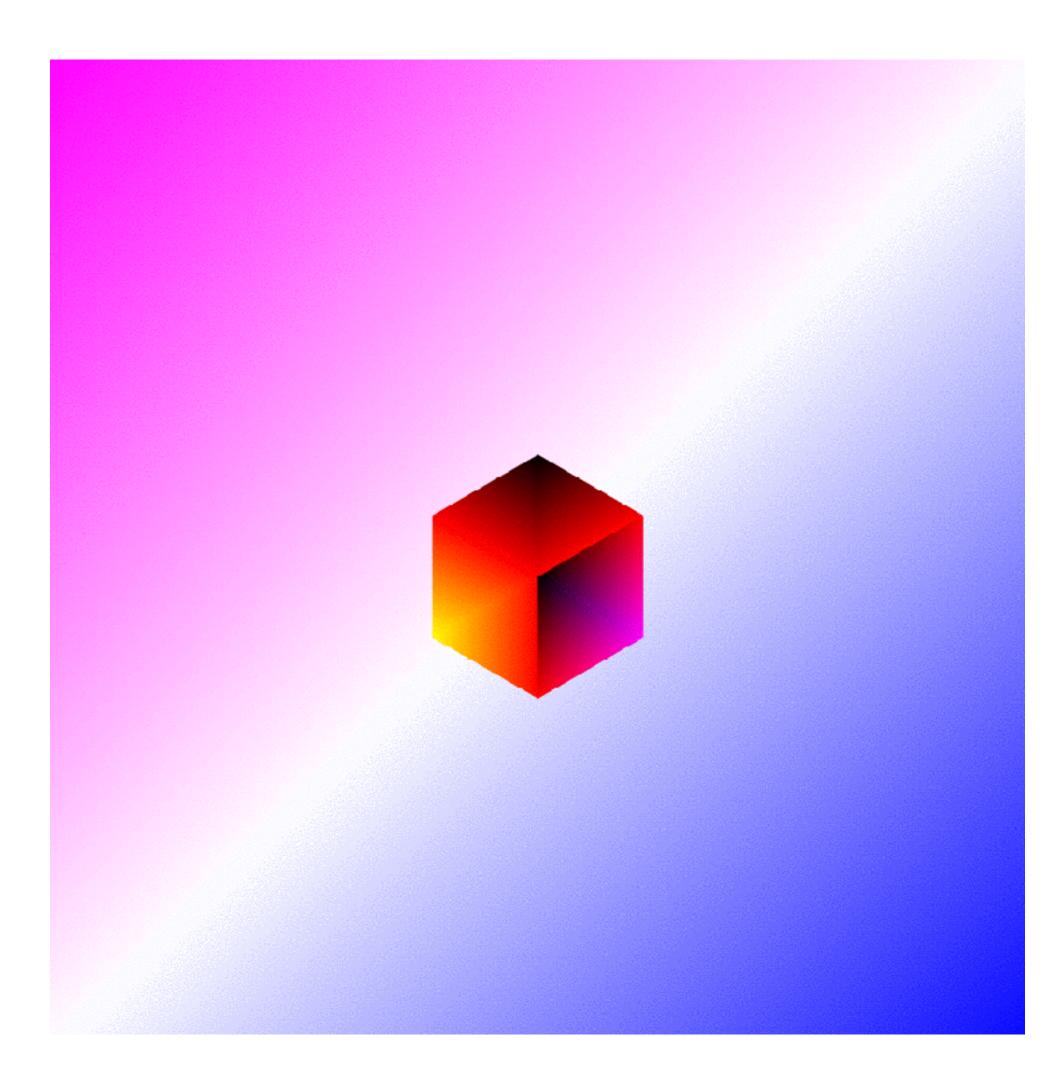
#### Jason Qiao

The happiest sad moment in film history.

#### Approach:

I created a triangle image generator in Python. It will iteratively place triangles and "jitter" them to find a local optimum, comparing our new image with the source image using RMSE. The triangles get smaller as more are placed so we can capture the details.

#### Art Competition #1 – 1st Place Winner



#### Ariel Hirschberg

"Cubeman's city is falling apart! Oh wait it's rebuilding itself. Perpetual motion?"

#### Approach:

- Python script outputs 74 SVG files describing a decent number of triangles.
- Project code (with a modified DrawRend class to match the desired 800x800 viewport) outputs 74 png files.
- Ezgif online stitches the pngs together into the final GIF.